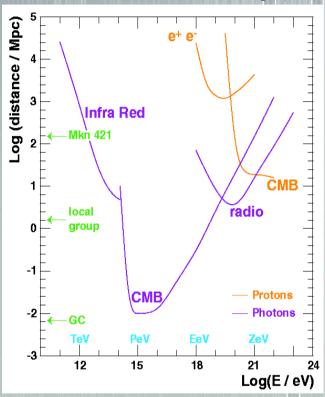
NEUTRINO ASTRONOMY

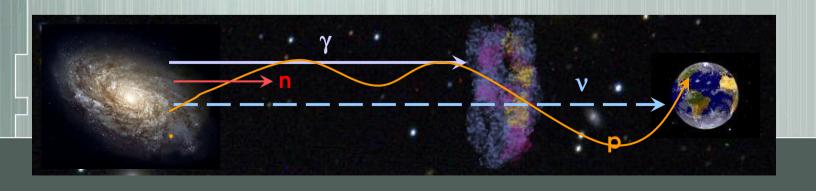
 Neutrino Astronomy is a quite recent and very promising experimental field.

Advantages:

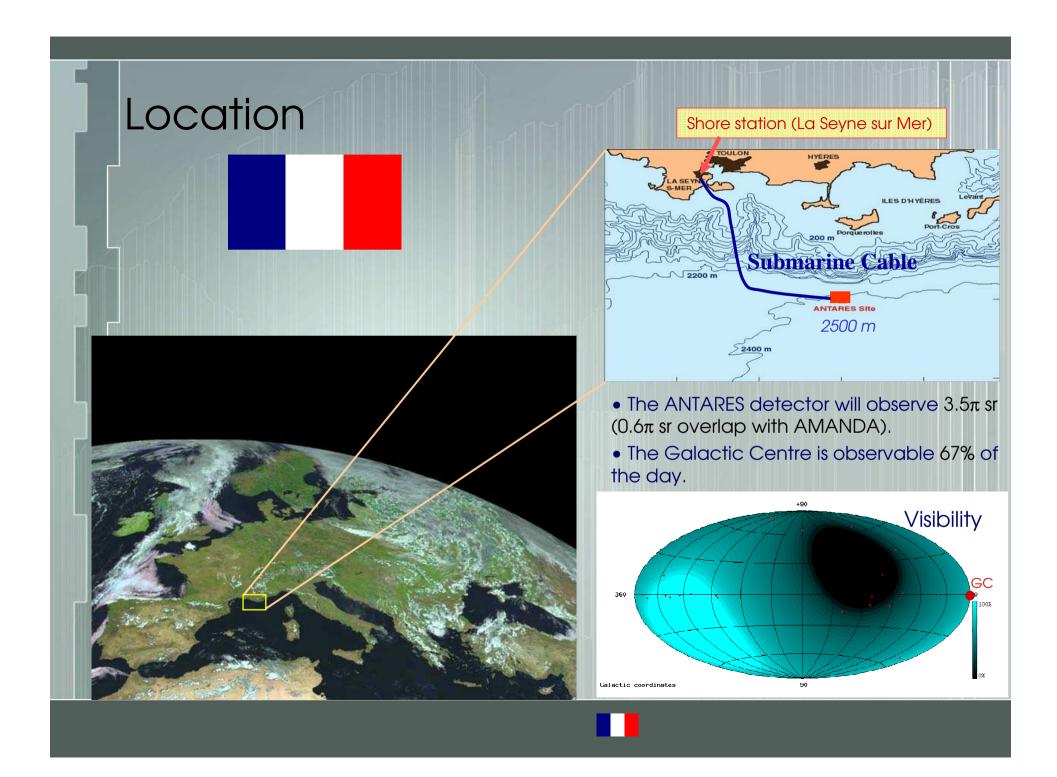
- Photons: interact with photons and matter
- Protons: interact with CMB and are deflected by magnetic fields
- Neutrons: are not stable
- <u>Drawback</u>: large detectors (~GTon) are needed.

Photon and **proton** mean free range path



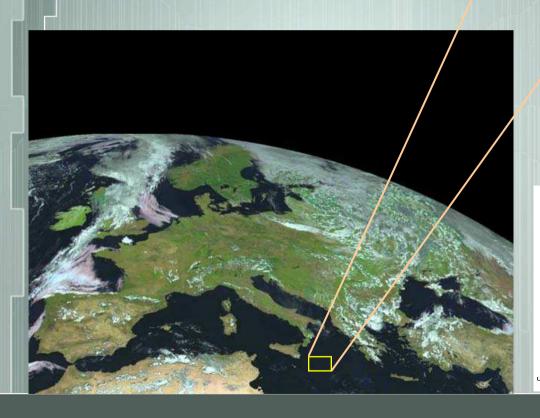


Scientific Scopes of Neutrino Telescopes $\Delta m^2 = 3.10^{-3} \text{ eV}^2$ GeV-100 ~MeV **Energy GeV-TeV** TeV-PeV PeV-EeV >EeV GeV **Astrophysical** Neutrino sources Neutralino AGNs, TD, Supernovae oscillation **Physics** (AGNs, GRBs, search GZK neutrinos MQs) Average **Almost Signature** Up-going Up-going **Up-going** Down-going increase in horizontal muons and muons muons tracks the PMT tracks cascades counting rate



Location

Shore station (Capo Passero)





- The NEMO detector will observe 3.5π sr $(0.6\pi$ sr overlap with AMANDA).
- The Galactic Centre is observable 67% of the day.

