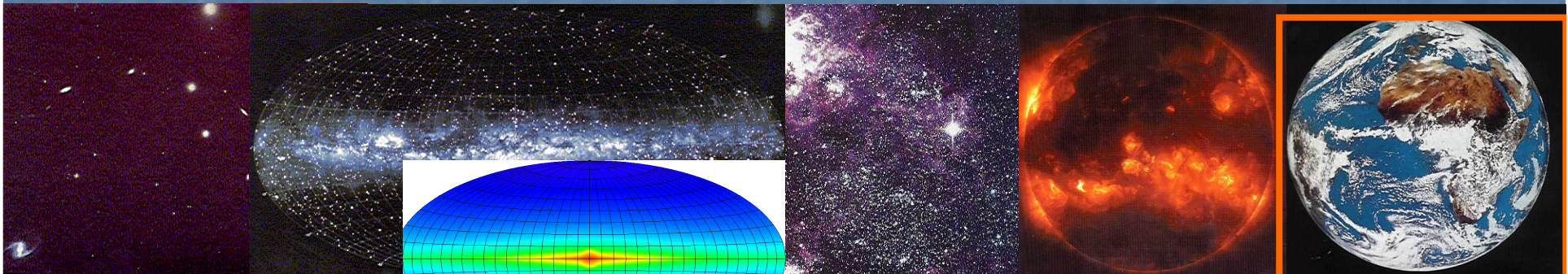


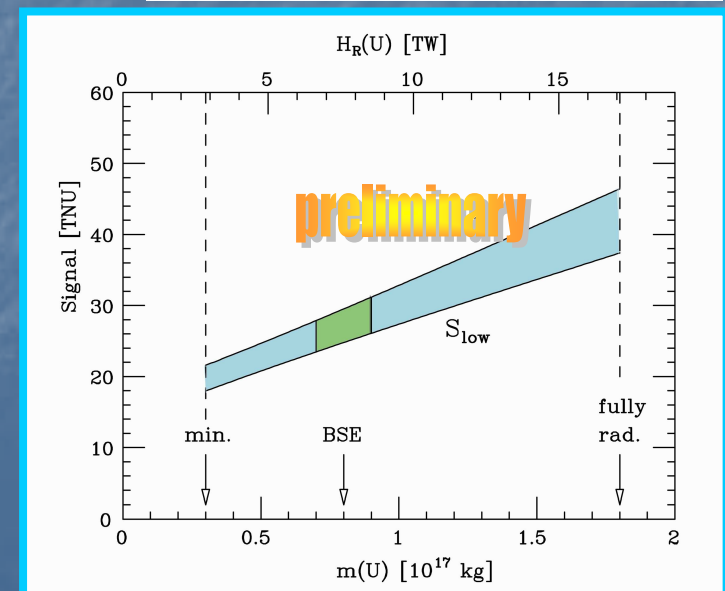
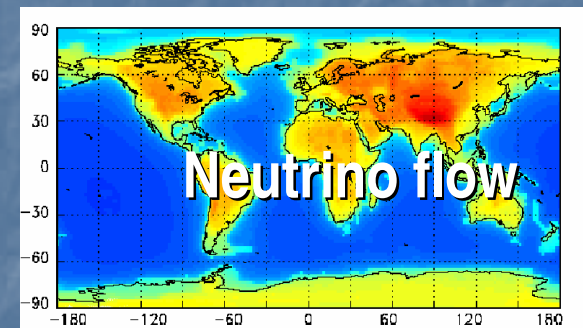
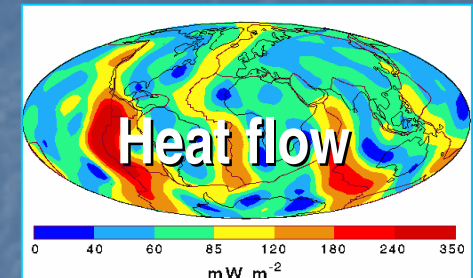
## On the shoulders of giants: a new era of neutrino physics ?

- We have still a lot to learn on neutrino properties (**M/ D**) and mass matrix however...
- Now we know the fate of neutrinos and **we can learn a lot from neutrinos.**



# Geo-Neutrinos : a new probe of Earth's interior

- What is the amount of U, Th and  $^{40}\text{K}$  in the Earth?
- Determine the radiogenic contribution to terrestrial heat flow
- Get information about the origin of the Earth.
- Test a fundamental geochemical paradigm: the Bulk Silicate Earth





# A lesson from Bruno Pontecorvo: from neutrons to neutrinos

**Neutron Well Logging - A New Geological Method Based on Nuclear Physics Oil and Gas Journal, 1941, vol.40, p.32-33.1942.**

•An application of Rome celebrated study on slow neutrons, **the neutron log** is an instrument sensitive to Hydrogen containing substances (=water and hydrocarbons), used for oil and water prospection.

•Now that we know the fate of neutrinos, we can learn a lot from neutrinos.

•The determination of the radiogenic contribution to Earth energetics is an important scientific question, possibly the first fruit we can get from neutrinos.

# Geo-neutrinos: anti-neutrinos from the Earth



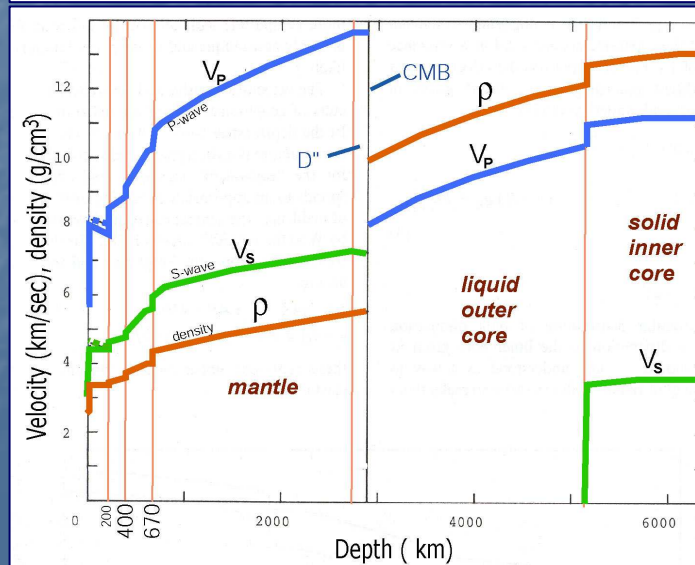
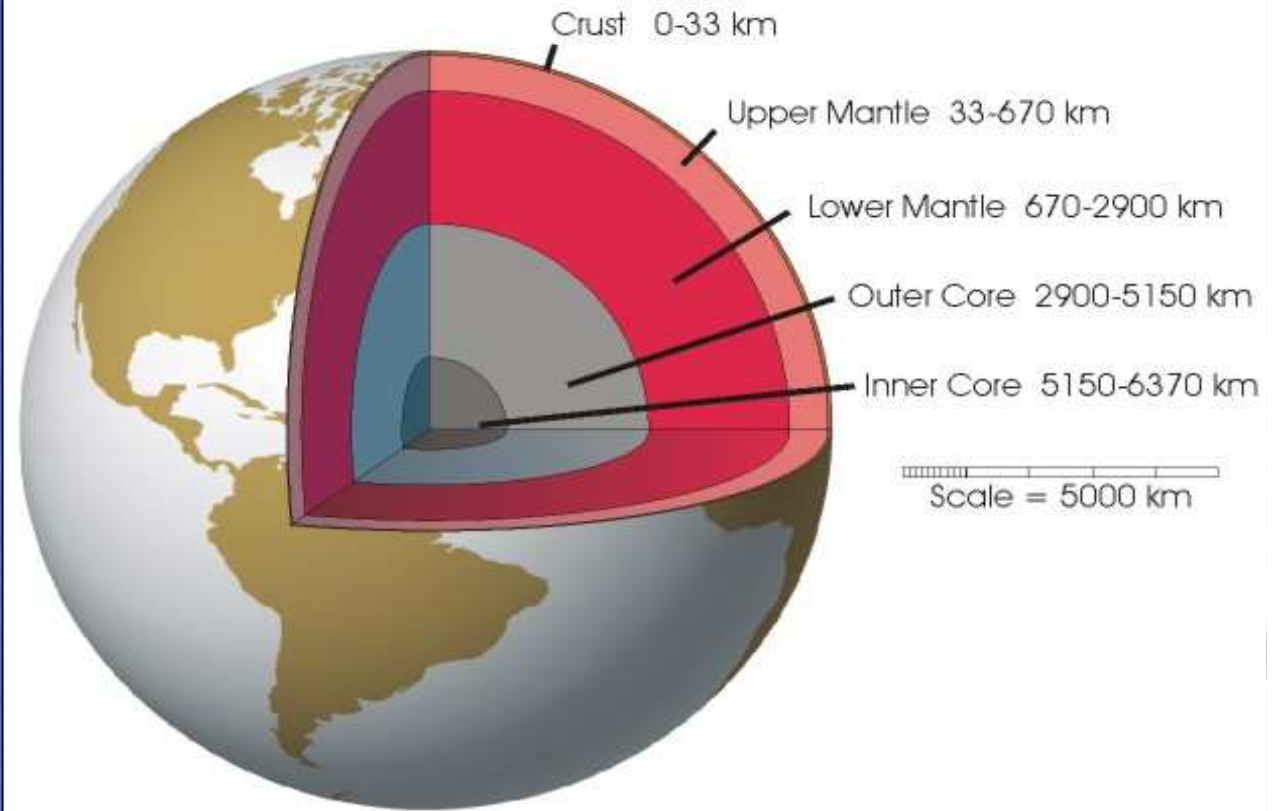
- Uranium, Thorium and Potassium in the Earth release heat together with anti-neutrinos, in a well fixed ratio:

Decay	Q [MeV]	$\tau_{1/2}$ [ $10^9$ yr]	$E_{max}$ [MeV]	$\epsilon_H$ [W/kg]	$\epsilon_{\bar{\nu}}$ [ $\text{kg}^{-1}\text{s}^{-1}$ ]
$^{238}\text{U} \rightarrow ^{206}\text{Pb} + 8^4\text{He} + 6e + 6\bar{\nu}$	51.7	4.47	3.26	$0.95 \cdot 10^{-4}$	$7.41 \cdot 10^7$
$^{232}\text{Th} \rightarrow ^{208}\text{Pb} + 6^4\text{He} + 4e + 4\bar{\nu}$	42.8	14.0	2.25	$0.27 \cdot 10^{-4}$	$1.63 \cdot 10^7$
$^{40}\text{K} \rightarrow ^{40}\text{Ca} + e + \bar{\nu}$	1.32	1.28	1.31	$0.36 \cdot 10^{-8}$	$2.69 \cdot 10^4$

- Earth emits (mainly) antineutrinos, Sun shines in neutrinos.
- Geo-neutrinos from U and Th (**not** from K) are above treshold for inverse  $\beta$  on protons:  $\bar{\nu} + p \rightarrow e^+ + n - 1.8\text{MeV}$
- Different components can be distinguished due to different energy spectra: anti- $\nu$  with highest energy are from Uranium

# Probes of the Earth's interior

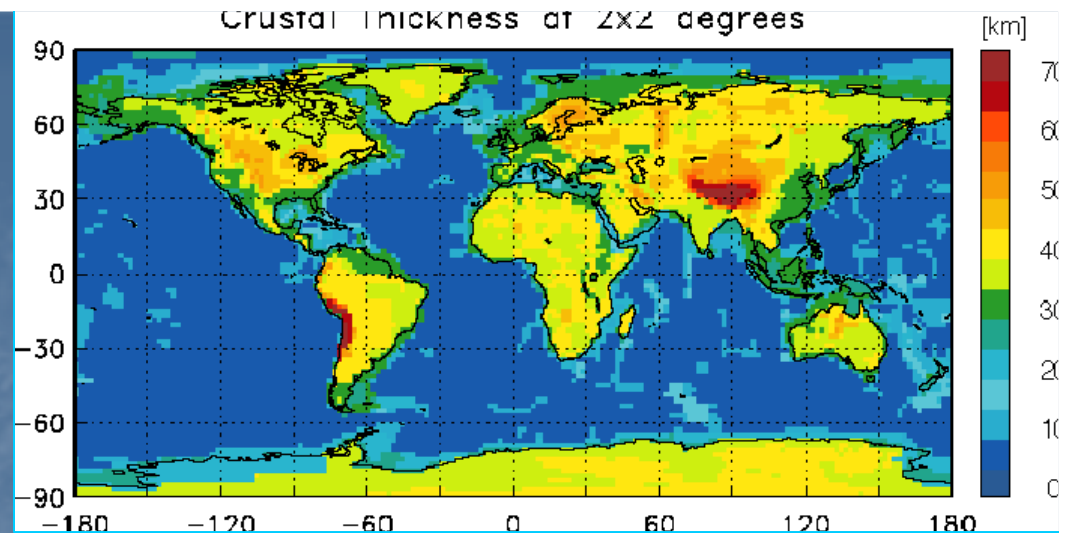
- Deepest hole is about 12 km.
- The crust (and the upper mantle only) are directly accessible to geochemical analysis.
- Seismology reconstructs density profile (not composition) throughout all earth.



• Geo-neutrinos can bring information about the chemical composition (U, Th and K) of the whole Earth.

# Uranium in the Earth: observational data on the crust

- Crust is the tiny envelope of the Earth, distinguished from the underlying mantle by a clear (Moho) seismic discontinuity.

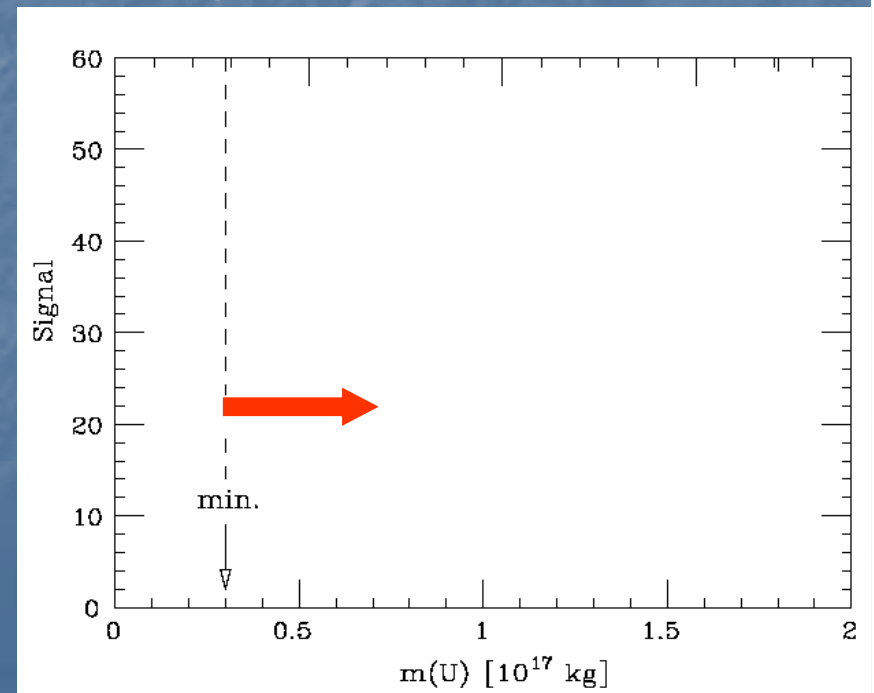


- Continental and oceanic crust have different origin and U abundance.

- By combining data on Uranium abundances from selected samples with geological maps of Earth's crust one concludes:

$$m_C(U) = (0.3-0.4)10^{17} \text{ kg}$$

- Most of the uncertainty from lower portion of the crust



# The amount of Uranium in the Earth: cosmo-chemical arguments

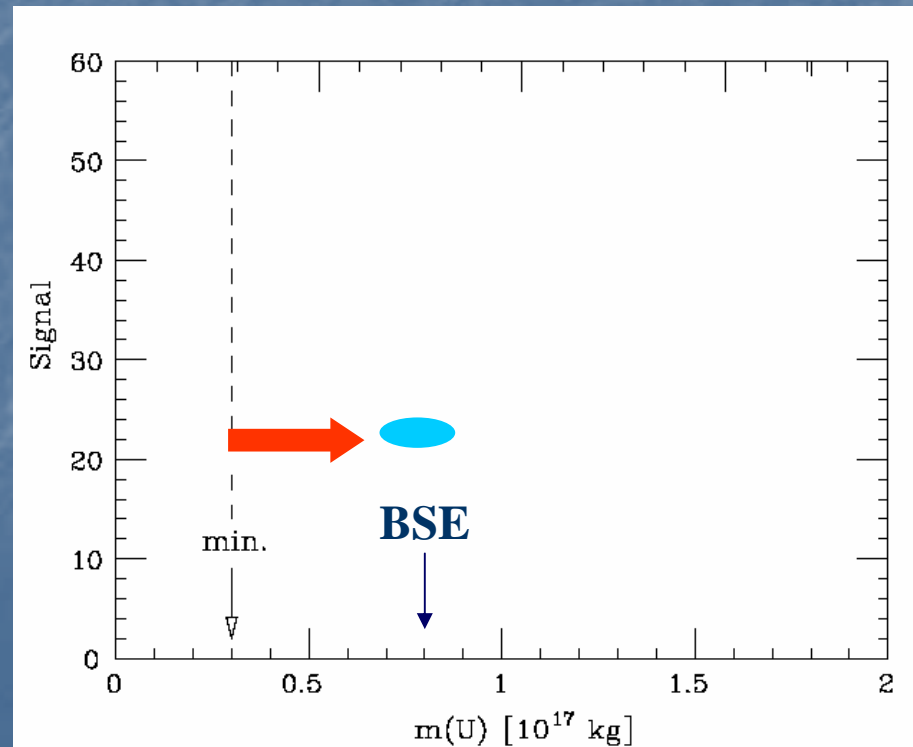


- The material form which Earth formed is generally believed to have same composition as CI-chondrites.

- By taking into account losses and fractionation in the initial Earth one builds the “Bulk Silicate Earth” (**BSE**), the standard geochemical paradigm which predicts:

$$m(\text{U}) = (0.7 - 0.9) \cdot 10^{17} \text{ kg}$$

- **Remark:** The BSE is grounded on solid geochemical + cosmochemical arguments, it provides a composition of the Earth in agreement with most observational data, however it lacks a direct observational test.

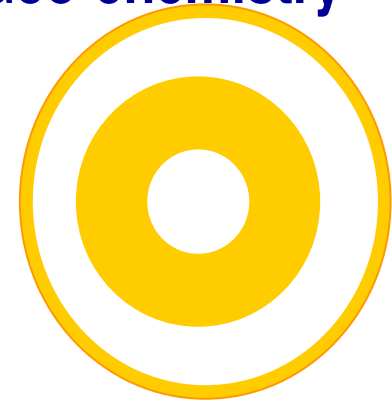


# Where is the rest of Uranium?

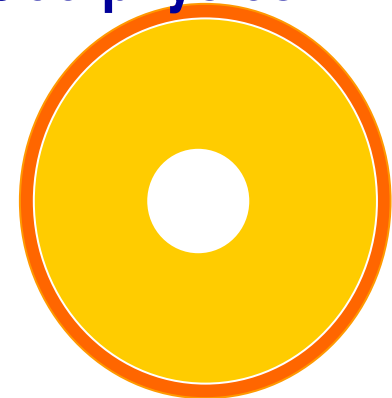
- According to BSE, crust contains about one half of the total Uranium amount
- Uranium is a lithophile element, believed (by geochemists) to be absent from the core.
- So the remaining half should be in the mantle:
  - A) according to geochemists, mainly in the lower part.
  - B) geophysics, indicating a globally homogeneous mantle, suggests a uniform distribution within the mantle.



## Geo-chemistry



## Geo-physics





# Heat released from the Earth

- The tiny flux of heat coming from the Earth ( $\Phi \approx 60 \text{ mW/m}^2$ ) when integrated over the Earth surface gives a total flow:

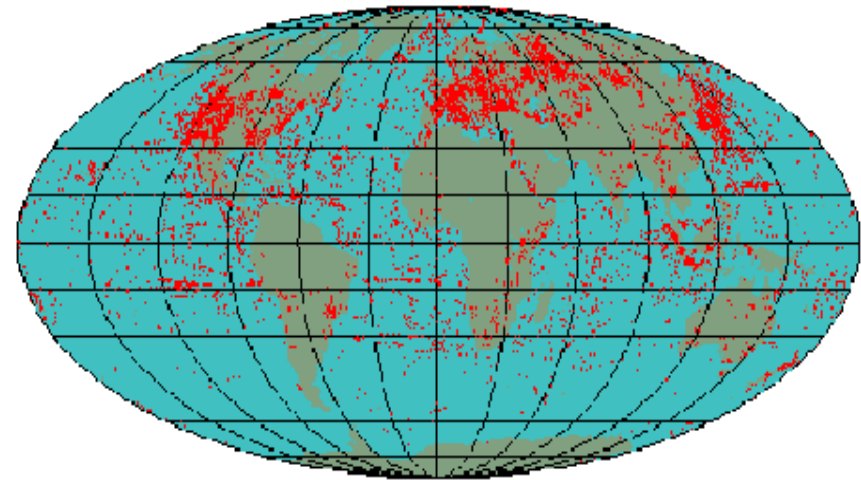
$$H_E = (30-45) \text{ TW}$$

- It is equivalent to  $10^4$  nuclear power plants.

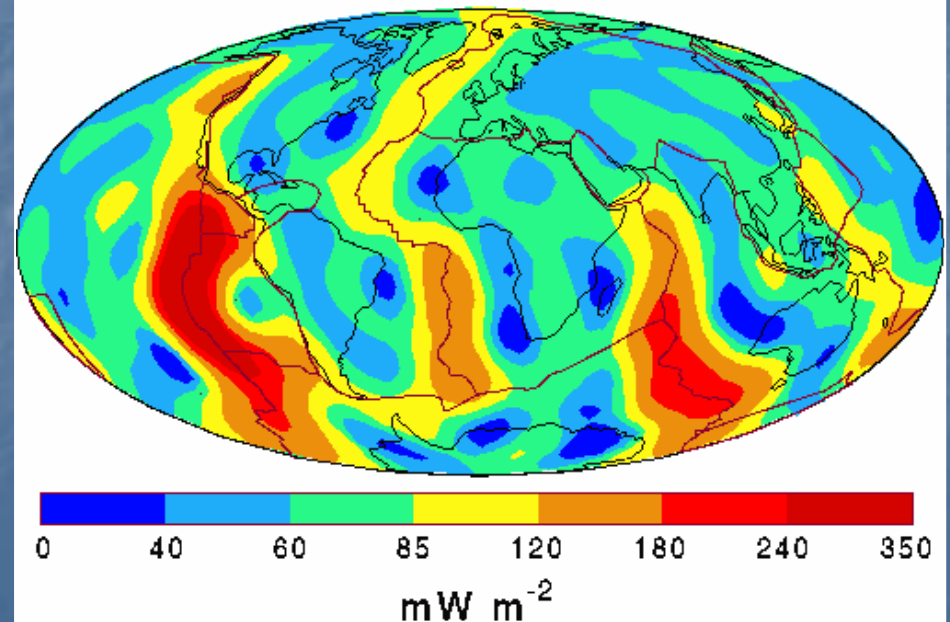
- Warning: the classical  $44 \pm 1 \text{ TW}$  (Pollack 93) recently revised to the “old”  $31 \pm 1 \text{ TW}$  (Hofmeister & Criss 04)

- What is its origin?

Heat Flow Sites



Heat Flow



# 2004



## Energetics of the Earth and the Missing Heat Source Mystery



[Don L. Anderson](#)

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[dla@gps.caltech.edu](mailto:dla@gps.caltech.edu)

**BSE**

Global heat flow estimates range from 30 to 44 TW ...  
Estimates of the radiogenic contribution, ... based on  
cosmochemical considerations, vary from 19 to 31 TW. Thus,  
there is either a good balance between current input and  
output, as was once believed ... or there is a serious missing  
heat source problem, up to a deficit of 25 TW...

• **Determination of the radiogenic component is important.**

# How much Uranium can be tolerated by Earth energetics?

- For each element there is a well fixed relationship between heat presently produced and its mass:

$$H_R = 9.5 m(U) + 2.7 m(\text{Th}) + 3.6 m(^{40}\text{K})$$

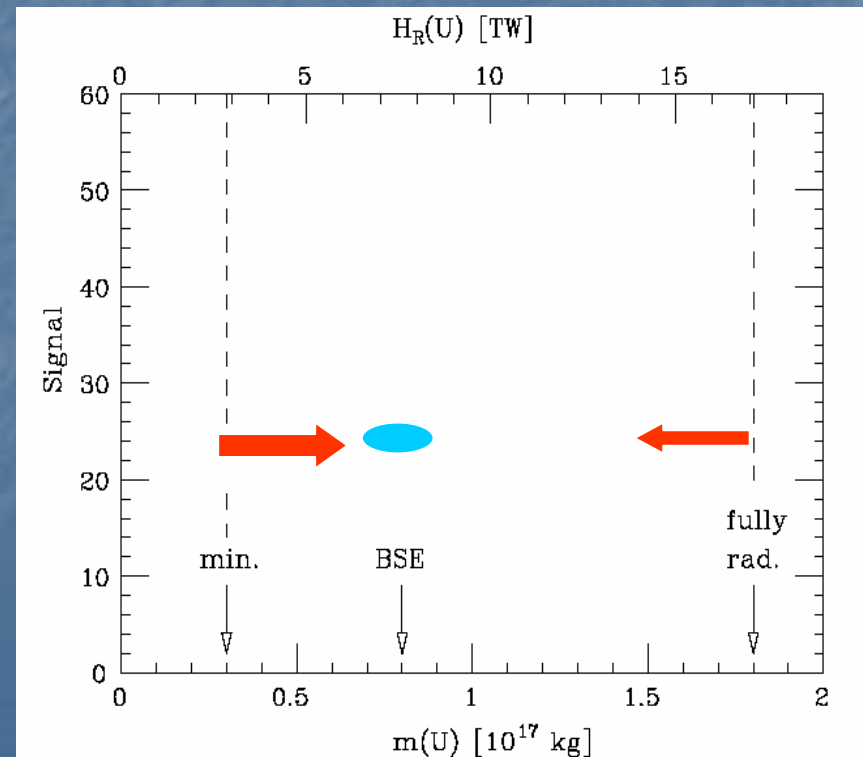
where units are TW and  $10^{17}\text{kg}$ .

- Since  $m(\text{Th}) : m(\text{U}) : m(^{40}\text{K}) = 4 : 1 : 1$

one has:  $H_R = 24 M(\text{U})$

- Present radiogenic heat production cannot exceed heat released from Earth:

$$m(\text{U}) < 1.8 \cdot 10^{17} \text{ kg}$$



# Order of magnitude estimate for the signal

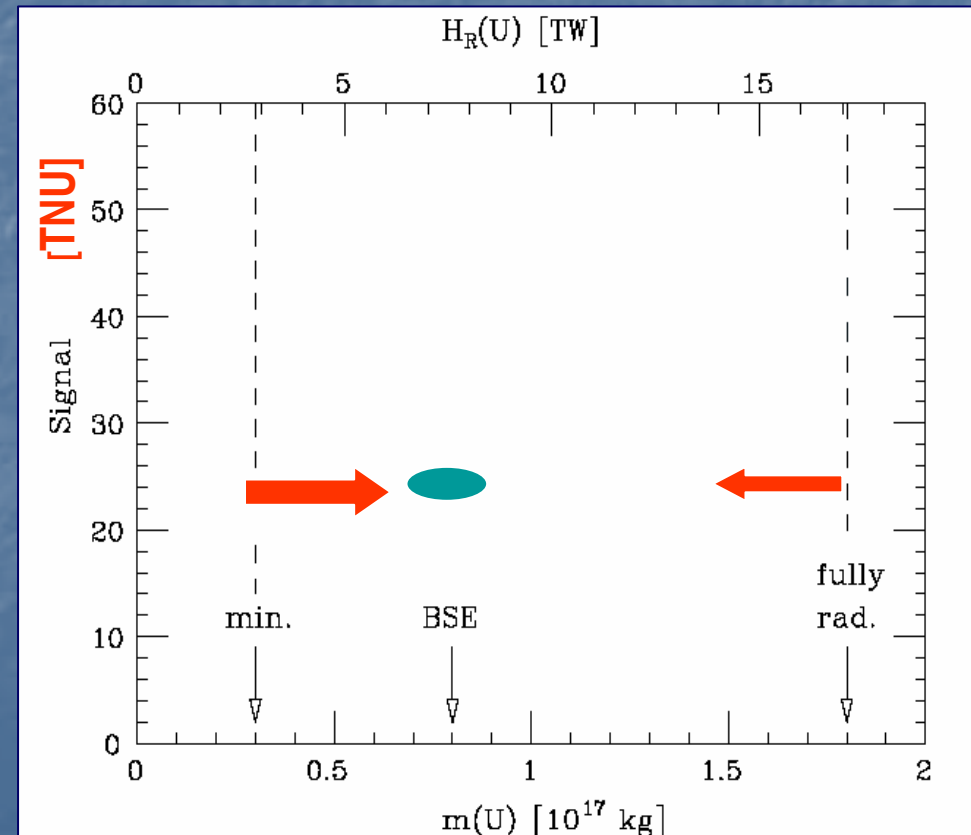
- From  $m(U)$  one immediately derives the geo-neutrino luminosity  $L$ , and an estimate for the flux  $\Phi \approx L/4\pi R_{\text{Earth}}^2$
- Fluxes are of order  $10^6 \nu \text{ cm}^{-2} \text{ s}^{-1}$ , same as  $^8\text{B}$ .
- From spectrum and cross section one gets the signal:

$$S = 13.2 \left( \frac{\Phi_{ar}}{10^6 \text{ cm}^{-2} \text{ s}^{-1}} \right) \left( \frac{N_p}{10^{32}} \right) \text{ yr}^{-1}$$

- Signal is expressed in Terrestrial Neutrino Units:

1 TNU = 1 event / (10<sup>32</sup> prot · yr)

(1kton LS contains 0.8 10<sup>32</sup> prot )



# The geo-neutrino signal and the Uranium mass: the strategy

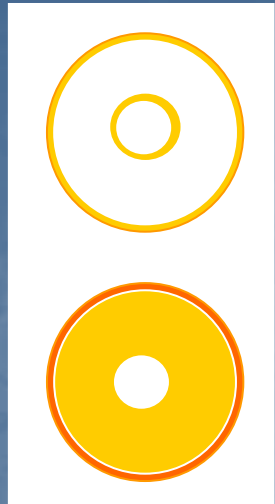
- Goal is in determining  $m(U)$  from geo- neutrino measurements.
- Signal will also depend on where detector is located:
- For  $m(U)=m_{BSE}$  we expect at Kamioka:

- $\frac{1}{2}$  of the signal from within 200 km
- This requires a detailed geochemical & geophysical study of the area.
- It is unsensitive to  $m(U)$

- The remaining  $\frac{1}{2}$  from the rest of the world.
- this is the part that brings information on  $m(U)$

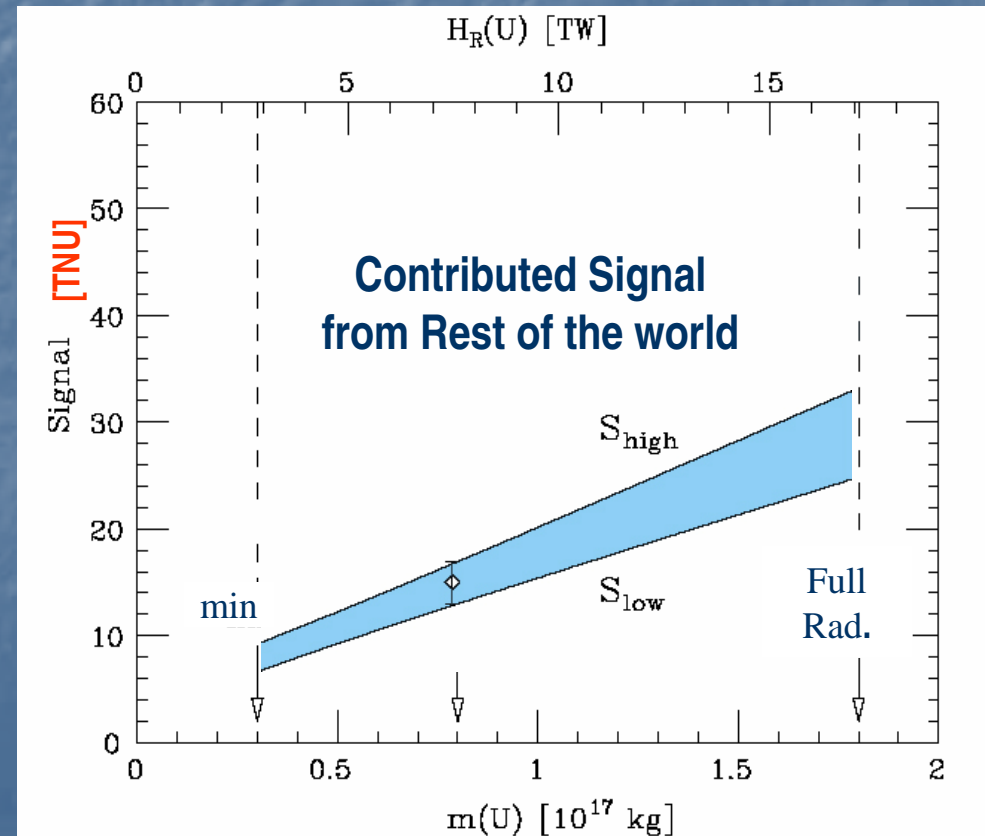
# The rest of the world.

- Signal depends on the value of Uranium mass and on its distribution inside Earth.
- For a fixed  $m(U)$ , the signal is maximal (minimal) when Uranium is as close (far) as possible to to detector:



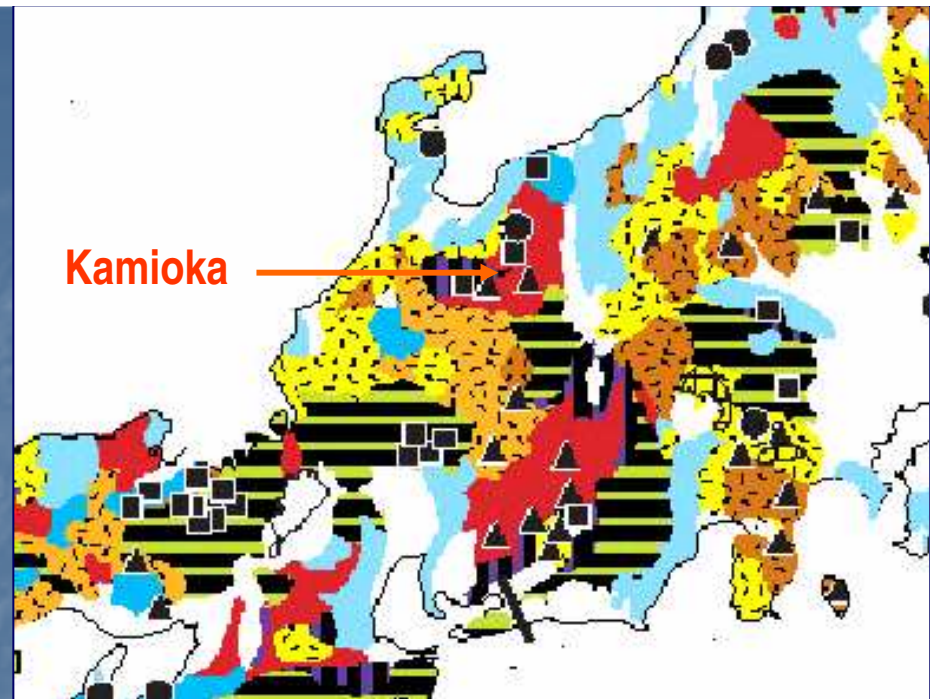
Signal	U in the Crust	U in the Mantle
Low	Poor	Retreated
High	Rich	Homog.

• Given  $m(U)$ , the signal from the rest of the world is fixed within  $\pm 10\%$



## The region near Kamioka

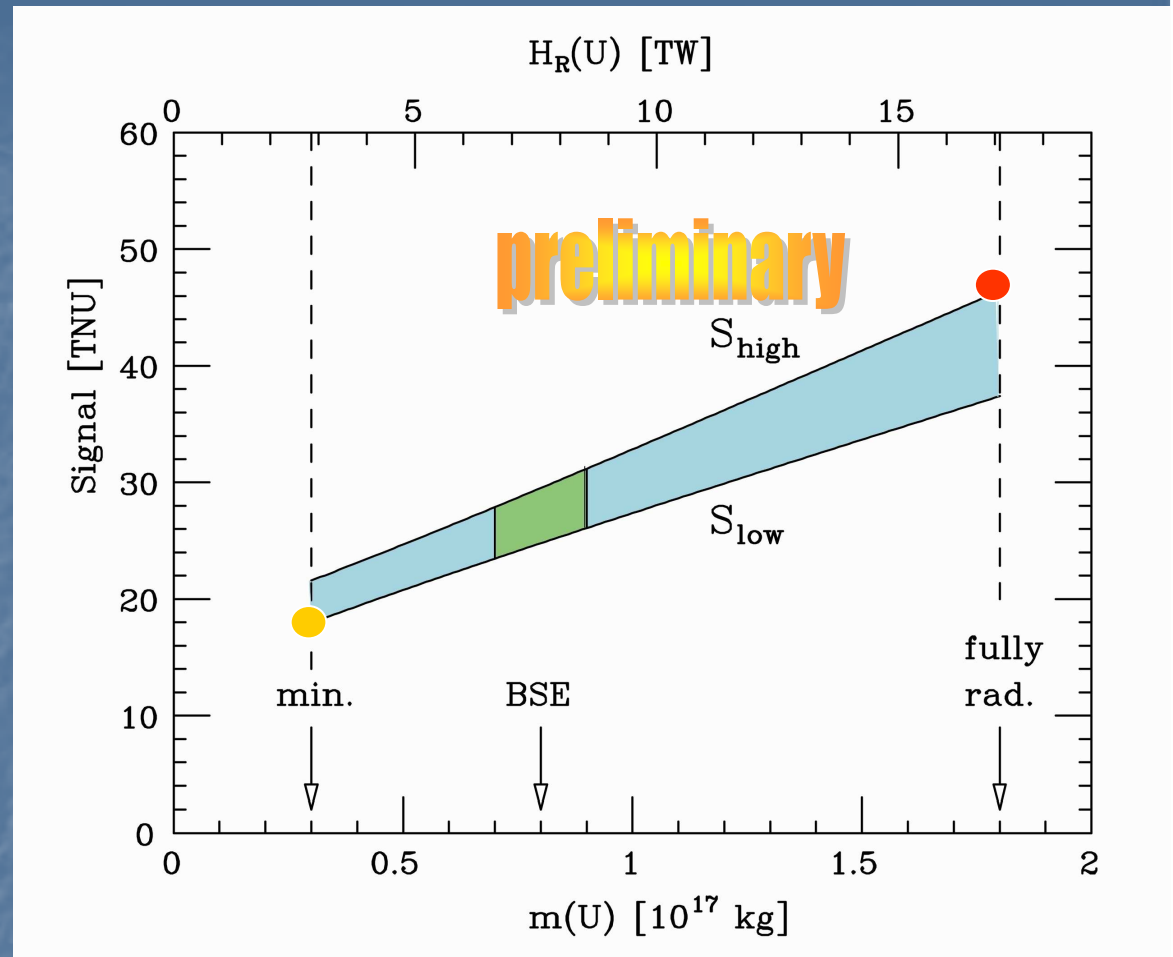
- Use a geochemical study of the Japan upper crust (scale  $\frac{1}{4} \times \frac{1}{4}$ ) and detailed measurements of crust depth.
- Use selected values for LC
- Take into account:
  - $(3\sigma)$  errors on sample activity measurements
  - Finite resolution of geochemical study
  - Uncertainty from the Japan sea crust characterization
  - Uncertainty from subducting plates below Japan
  - Uncertainty of seismic measurements



- In this way the accuracy on the local contribution can be matched with the uncertainty of the global estimate.

# Geo-neutrino signal at Kamioka and Uranium mass in the Earth

- 1) Uranium measured in the crust implies a signal of at least **18 TNU**
- 2) Earth energetics implies the signal does not exceed **46 TNU**
- 3) **BSE** predicts a signal between **23 and 31 TNU**



Geo-neutrino detection can provide a direct test of BSE prediction.



# Looking forward to new data

- KamLAND already provided a first glimpse  $S(U+Th)=(82\pm 52_{\text{stat.}})$  TNU

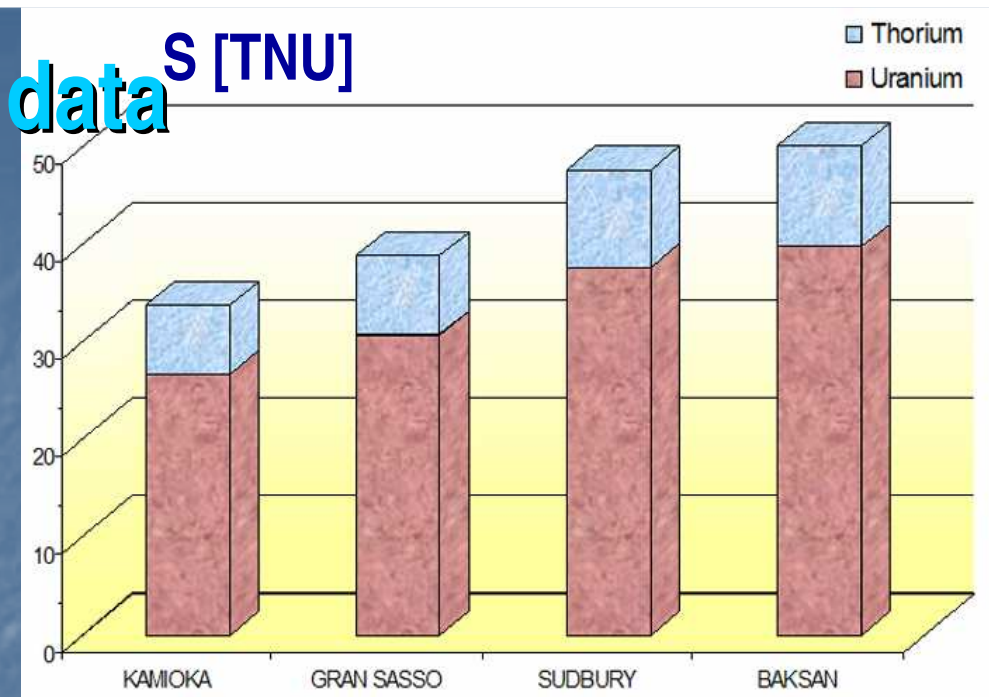
- KamLAND is analyzing data for geo-neutrinos now...

- Need to subtract reactor events, may be 10 times as many as geo-neutrino events.

- Borexino at Gran Sasso will have smaller mass but better geo/reactor .

- At Baksan Mikaelyan et al. are considering 1Kton detector, again far from nuclear reactors.

- LENA in Finland envisages a 30Kton LS detector



- At SNO there are plans of moving to liquid scintillator after physics  $D_2O$  is completed. With low reactor background, well in the middle of Canadian shield (an “easy geological situation) it will have excellent opportunities.

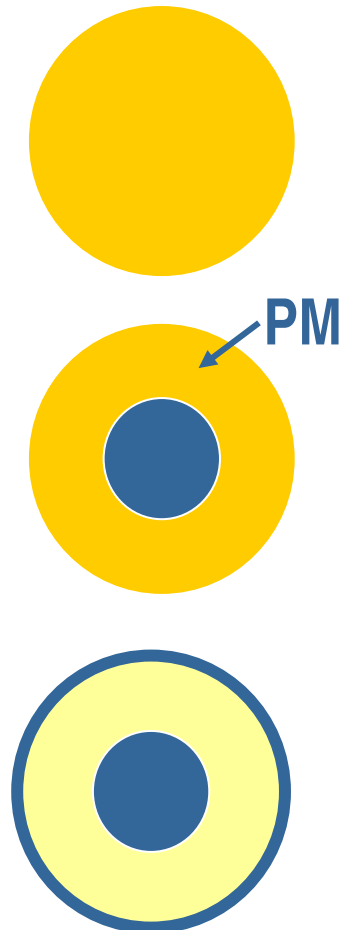
*“Se son rose, fioriranno...”*

# Appendix

# The canonical Bulk Silicate Earth paradigm



- CI chondritic meteorites are considered as representative of the primitive material of the solar system.
- Earth's global composition is generally estimated from that of CI by using geochemical arguments, which account for loss and fractionation during planet formation.
- In this way the Bulk Silicate Earth (BSE) model is built.
- It describes the “primitive mantle” i.e.:
  - subsequent to core formation.
  - prior to the differentiation between crust and mantle
- It is assumed to describe the present crust plus mantle.
- It is a fundamental geochemical paradigm, consistent with most observations. It should be tested.



# U, Th and K according to BSE

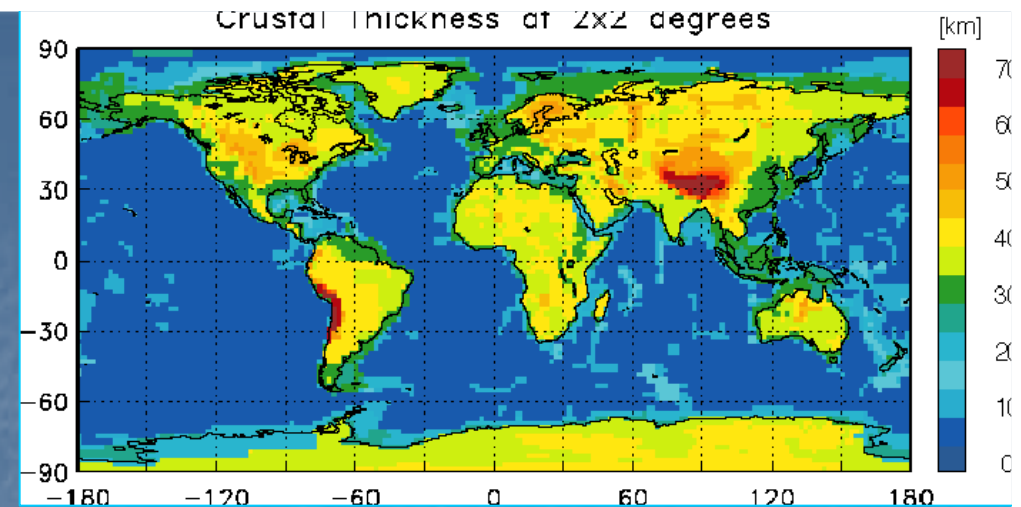
- Global masses of U, Th and K are estimated with accuracy of  $\pm 15\%$
- Radiogenic Heat and neutrino Luminosity can be immediately calculated:

	<b>M(<math>10^{17}</math>kg)</b>	<b>H<sub>R</sub>(TW)</b>	<b>L<sub>ν</sub>(<math>10^{24}</math>/s)</b>
<b>U</b>	<b>0.8</b>	<b>7.6</b>	<b>5.9</b>
<b>Th</b>	<b>3.1</b>	<b>8.5</b>	<b>5.0</b>
<b><sup>40</sup>K</b>	<b>0.8</b>	<b>3.3</b>	<b>21.6</b>

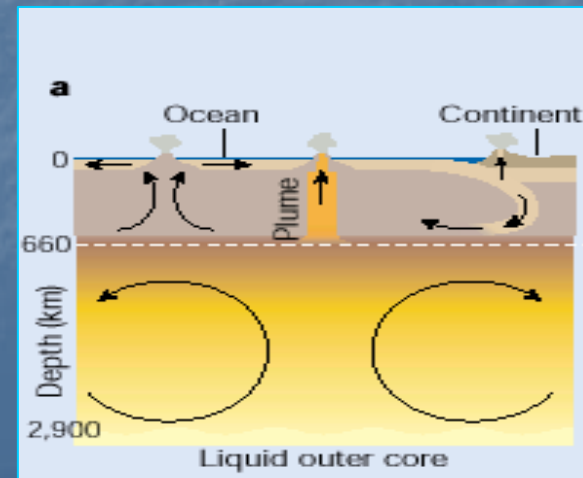
- Amounts U, Th and K inferred for the mantle are comparable to those observed in the crust
- Total radiogenic heat production (19 TW) is about  $\frac{1}{2}$  of observed heat flow, with comparable contribution from U and Th.
- Neutrino luminosity is dominated by K. Th and U give comparable contributions.

# From luminosity to fluxes

- Anti neutrino fluxes are of the order  $\Phi \approx L_\nu / S_{\text{Earth}} \approx 10^6 \text{ cm}^{-2} \text{ s}^{-1}$  [as for solar B-neutrinos].
- The flux at a specific site can be calculated from total amounts of radioactive nuclei and their distribution.
- The crust contribution can be estimated by using geological maps of Earth crust (which distinguish CC from OC and also distinguish several layers in the CC).

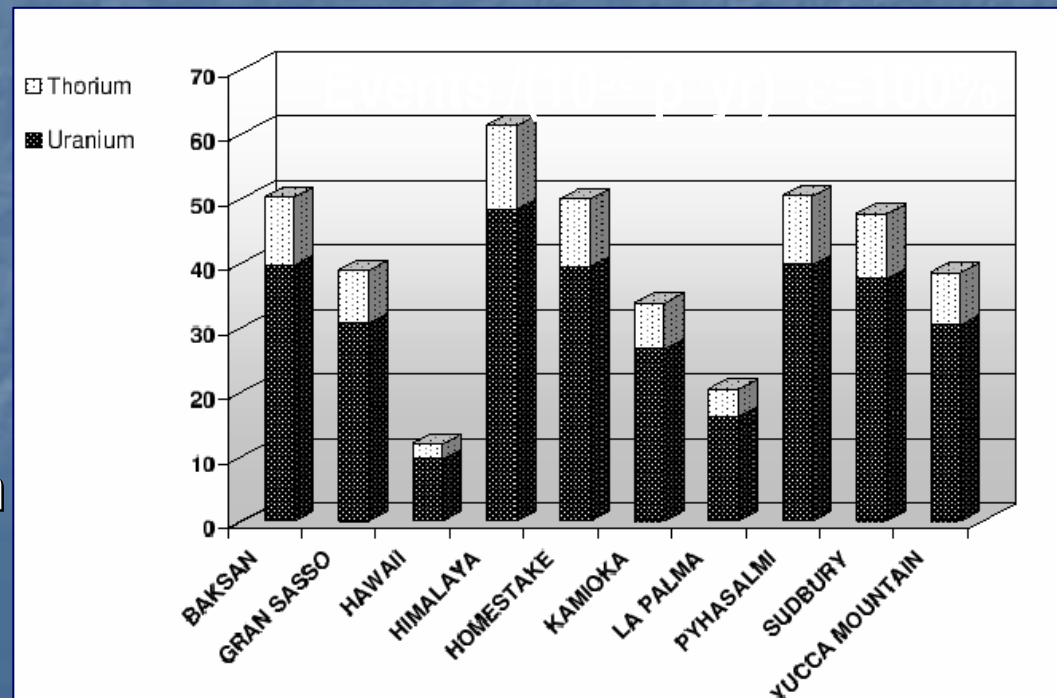
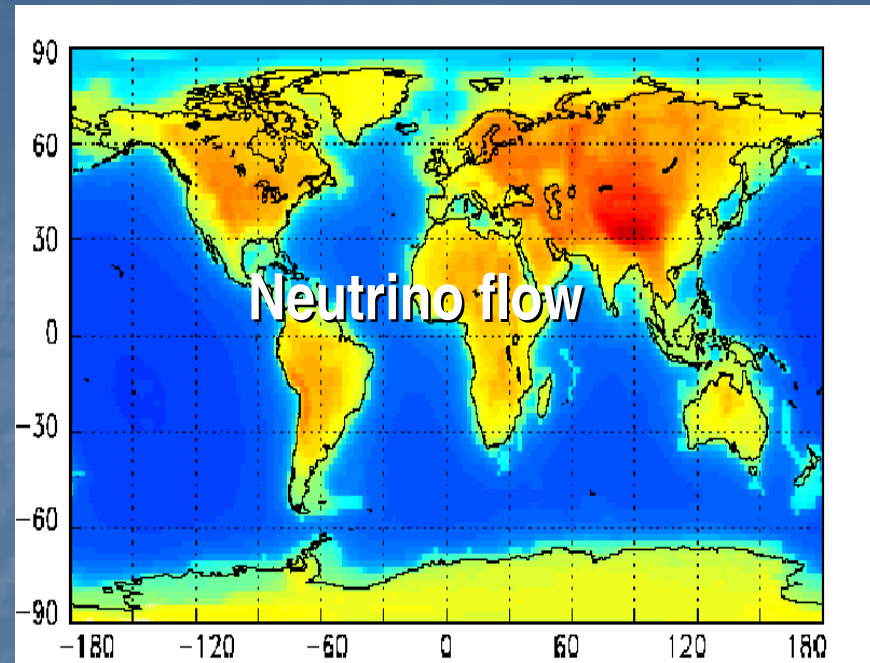


- The geochemist's mantle model is layered, the upper part being impoverished, abundance in the lower part being chosen so as to satisfy BSE mass balance.



# A reference BSE geo-neutrino model\*

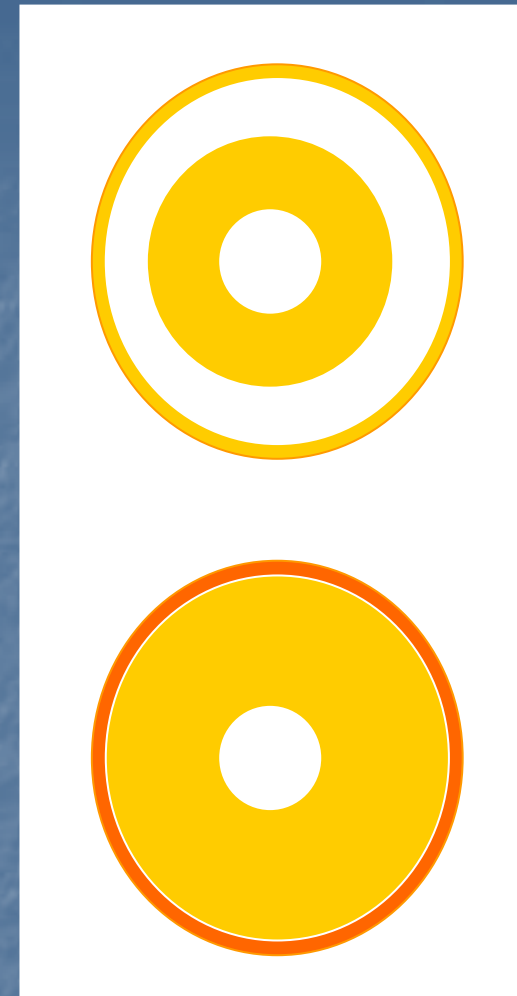
- Event yields from U and Th over the globe have been calculated by using:
  - observational data for Crust and UM
  - the BSE constraint for LM
  - best fit  $\nu$ -oscillation parameters
- Predicted events are about 30 per kiloton-yr, depending on location.
- $\frac{3}{4}$  originate from U,  $\frac{1}{4}$  from Th decay chains



\*Mantovani et al PRD-2003

# Testing the Bulk Silicate Earth with geo-neutrinos\*

- BSE fixes the total U mass ( to  $\pm 15\%$ )
- **The minimal (maximal) flux is obtained by putting the sources as far (as close) as possible.**
- The predicted flux contribution from distant sources in the crust and in the mantle is thus fixed within  $\pm 20\%$ .
- A detailed investigation of the region near the detector has to be performed, for reducing the uncertainty from fluctuations of the local abundances.
- A five-kton detector operating over four years at a site relatively far from nuclear power plants can measure the geo-neutrino signal with 5% accuracy



**It will provide a direct test of a fundamental geochemical paradigm**

# A word of caution

- CI based Bulk Silicate Earth (BSE) is the standard model of geochemists and its geo-neutrino predictions are rather well defined. It does not mean they are correct.
- Geo-neutrinos offer a probe for testing these predictions.
- Alternative models can be envisaged.
- A 40 TW (fully) radiogenic model ( with  $^{40}\text{K}:\text{U}:\text{Th}=1:1:4$ ) at 40 TW is not excluded by observational data.
- It needs  $M(\text{U}, \text{Th}, \text{K})=2x M_{\text{BSE}}(\text{U}, \text{Th}, \text{K})$ , most being hidden in LM

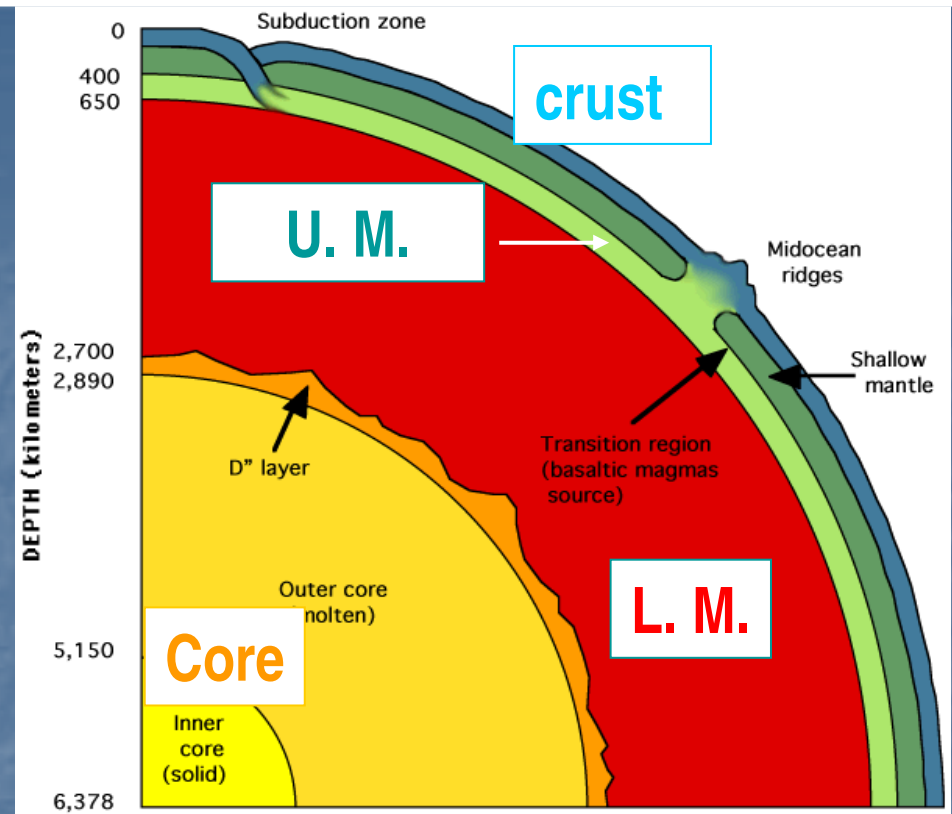
• Experiments should be designed so as to provide discrimination between BSE and FUL-RAD

	Events $/(10^{32} \text{ p} \cdot \text{yr}) \quad \epsilon=100\%$			
	Hawaii	Kam	GS	Himalay a
<b>BSE</b>	<b>12</b>	<b>33</b>	<b>39</b>	<b>62</b>
<b>Ful-Rad</b>	<b>27</b>	<b>53</b>	<b>58</b>	<b>85</b>



# Where are U, Th and K?

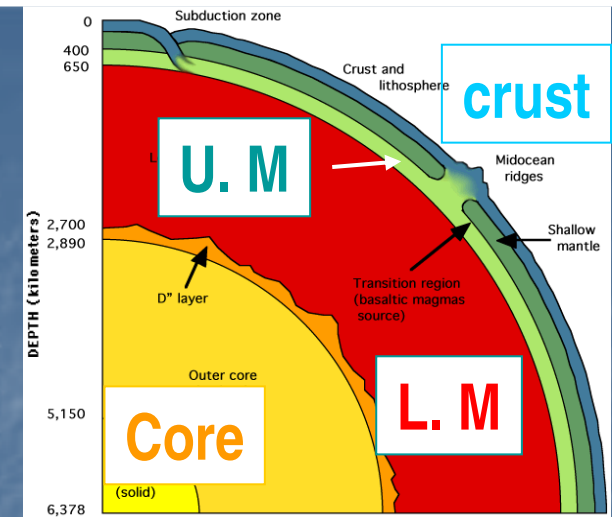
- The crust (and the upper mantle only) are directly accessible to geochemical analysis.
- U, K and Th are “lithophile”, so they accumulate in the (continental) crust.
- U in the crust is:  
 $M_c(U) \approx (0.3-0.4)10^{17} \text{Kg}$ .
- The  $\approx 30 \text{ Km}$  crust should contain roughly as much as the  $\approx 3000 \text{ km}$  deep mantle.
- Concerning other elements:  
 $\text{Th}/\text{U} \approx 4^*$  and  $^{40}\text{K}/\text{U} \approx 1$



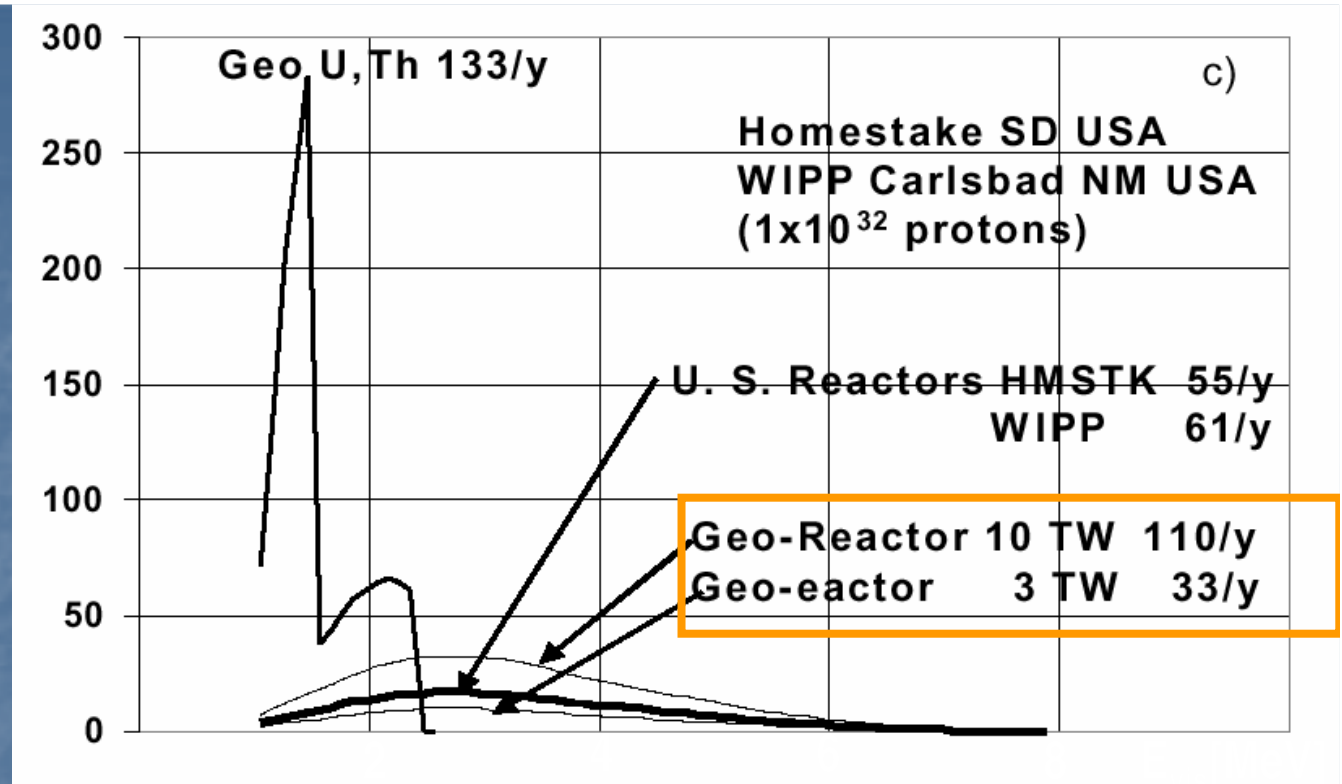
- For the lower mantle essentially no direct information: one relies on data from meteorites through geo-(cosmo)-chemical (BSE) model...
- According to geochemistry, no U, Th and K should be present in the core.

# Un-orthodox models: Potassium in the core?

- Earth looks depleted by a factor of seven with respect to CI meteorites.
- It has been suggested that missing Potassium might have been buried in the Earth core (although lithophile elements are not expected there).
- It could provide the energy source of the terrestrial magnetic field and a huge contribution to Earth energetics  $H_r(K) = 3.3 \times 7 = 23$  TW, solving the missing heat problem.
- The flux of Anti- $\nu$  from  $^{40}\text{K}$  at KamLAND would be  $10^8 \text{cm}^{-2}\text{s}^{-1}$ , but they are below threshold for inverse  $\beta$ .
- Indirectly, one can learn on K from U and Th geo-neutrinos: if U and Th are found to satisfy energy balance, no place is left for  $^{40}\text{K}$ .



# Heretical models: a nuclear reactor in the core?



- Herndon proposed that a large fraction of Uranium has been collected at the center of the Earth, forming a natural 3-6 TW (breeder) reactor.
- Fission should provide the energy source for mag. field, a contribution to missing heat, and the source of “high”  $^3\text{He}/^4\text{He}$  flow from Earth.
- Raghavan has considered possible detection by means of “reactor type antineutrinos”: a 1Kton detector in US can reach  $3\sigma$  in one year.
- Time dependence of man made reactor signal could be exploited.

# KAMLAND: a first important glimpse

- From six months data ( $0.14 \cdot 10^{32}$  p·yr) the KamLAND best fit is

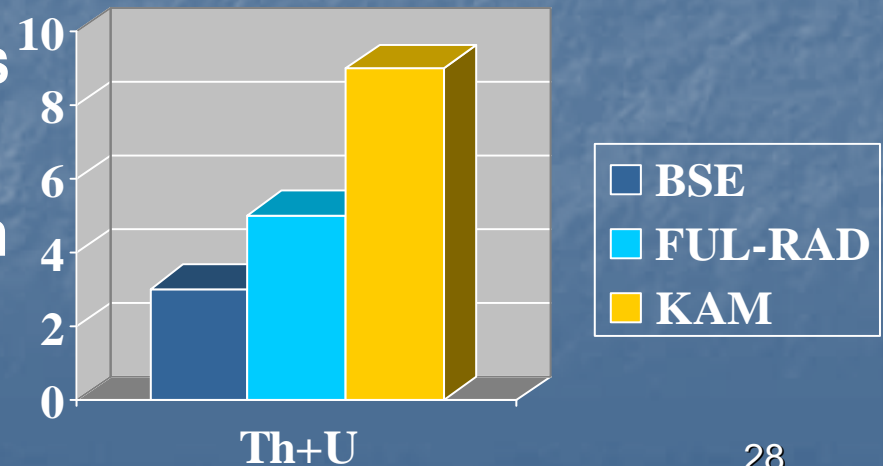
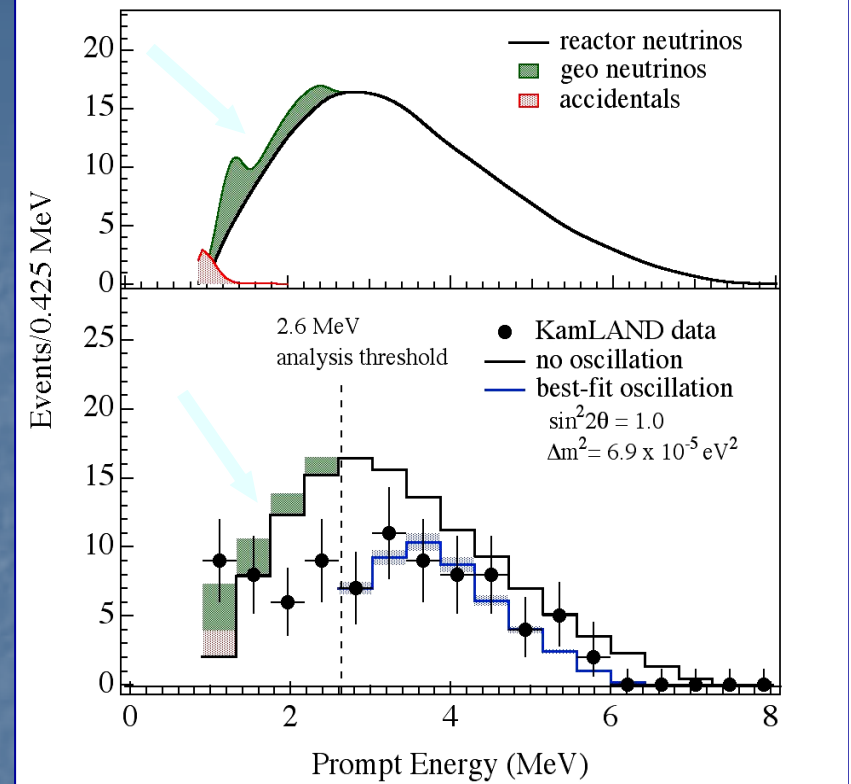
$$N(\text{U})=4 \text{ and } N(\text{Th})=5$$

- This results from 32 counts with P.E. < 2.6 MeV (20 attributed to reactor and 3 to B.G.) .

$$N(\text{Th+U}) = 9 \pm \sqrt{(\text{Counts})} = 9 \pm 6^*$$

- The error\* is dominated by fluctuations of reactor counts.
- The result is essentially consistent with any model,  $H_r=(0-100 \text{ TW})$ .
- Wait and see...

\* our estimate





## A few references\*



G.Eder, Nuc. Phys. 1966

G Marx Czech J. Phys. 1969, PR '81

Krauss Glashow, Schramm, Nature '84

Kobayashi Fukao Geoph. Res. Lett '91

Raghavan Schoenert Suzuki PRL '98

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Fiorentini et al PL 2002

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Mitsui ICRC 2003

Miramonti 2003

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McKeown Vogel, 2004

Fields, Hochmuth 2004

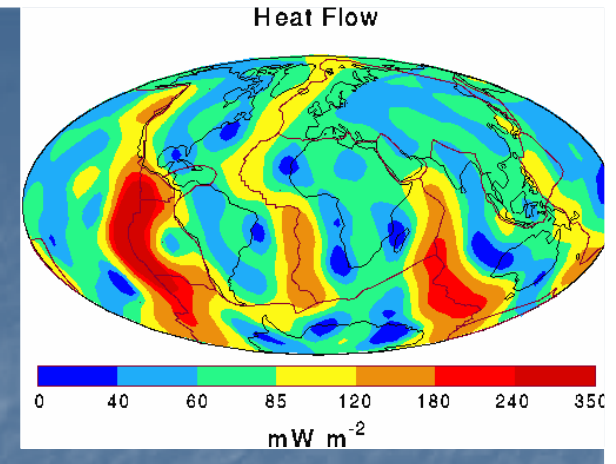
Fogli et al 2004

- Geo-neutrinos were introduced by G Eder and first discussed by G Marx
- More refs in the last 2 years than in previous 30.
- Most in the list are theoreticians, experimentalists added recently.

\*Apologize for missing refs.<sup>29</sup>

# What is the source of terrestrial heat?

J Verhoogen, in “Energetics of Earth” (1980)

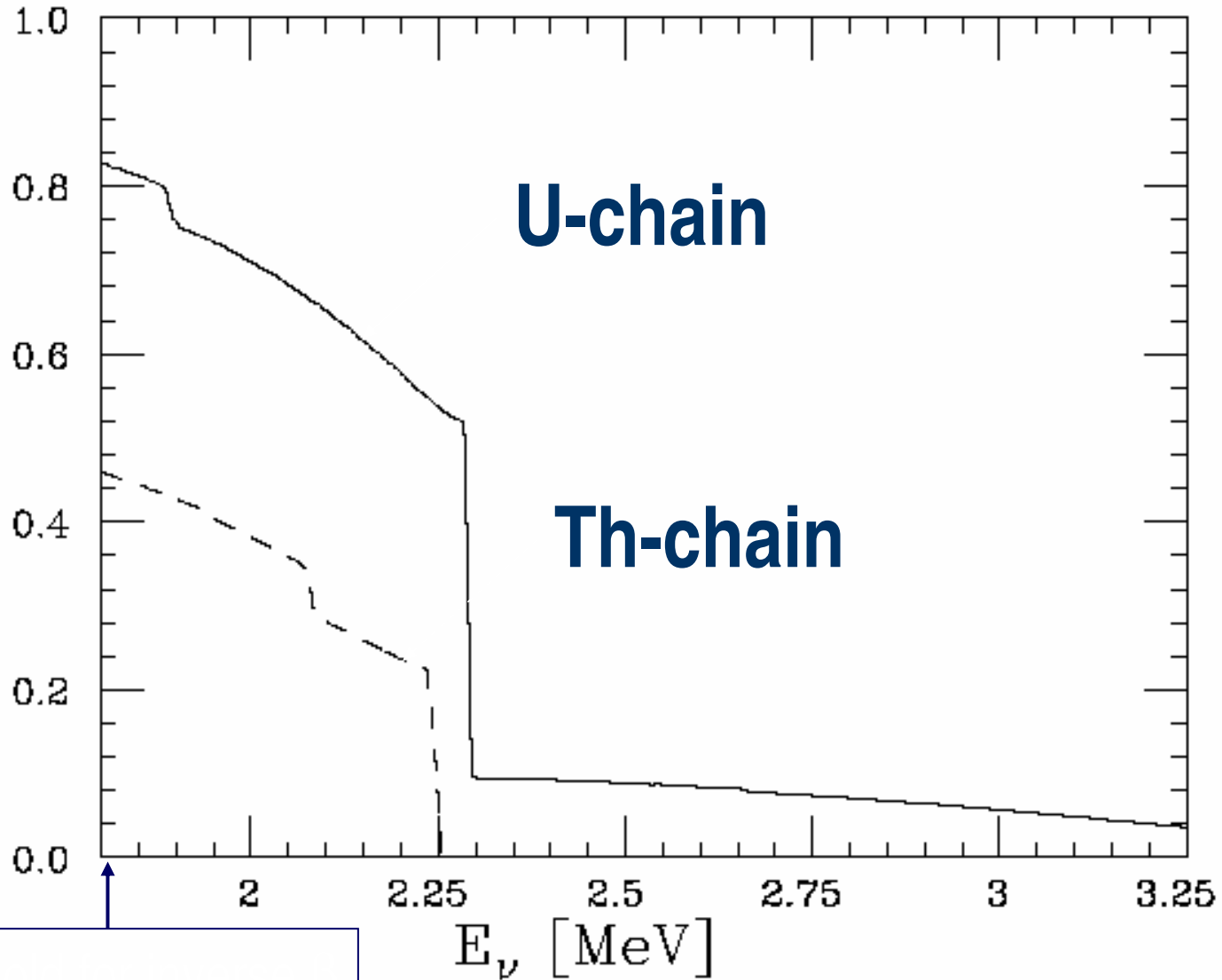


- “...What emerges from this morass of fragmentary and uncertain data is that **radioactivity** itself could possibly account for at least **60 per cent if not 100 per cent** of the Earth’s heat output”.
- “If one adds the greater rate of radiogenic heat production in the past, possible release of gravitational energy (original heat, separation of the core...) tidal friction ... and possible meteoritic impact ... the **total supply** of energy may seem **embarrassingly large...**”

• Determination of the radiogenic component is important.

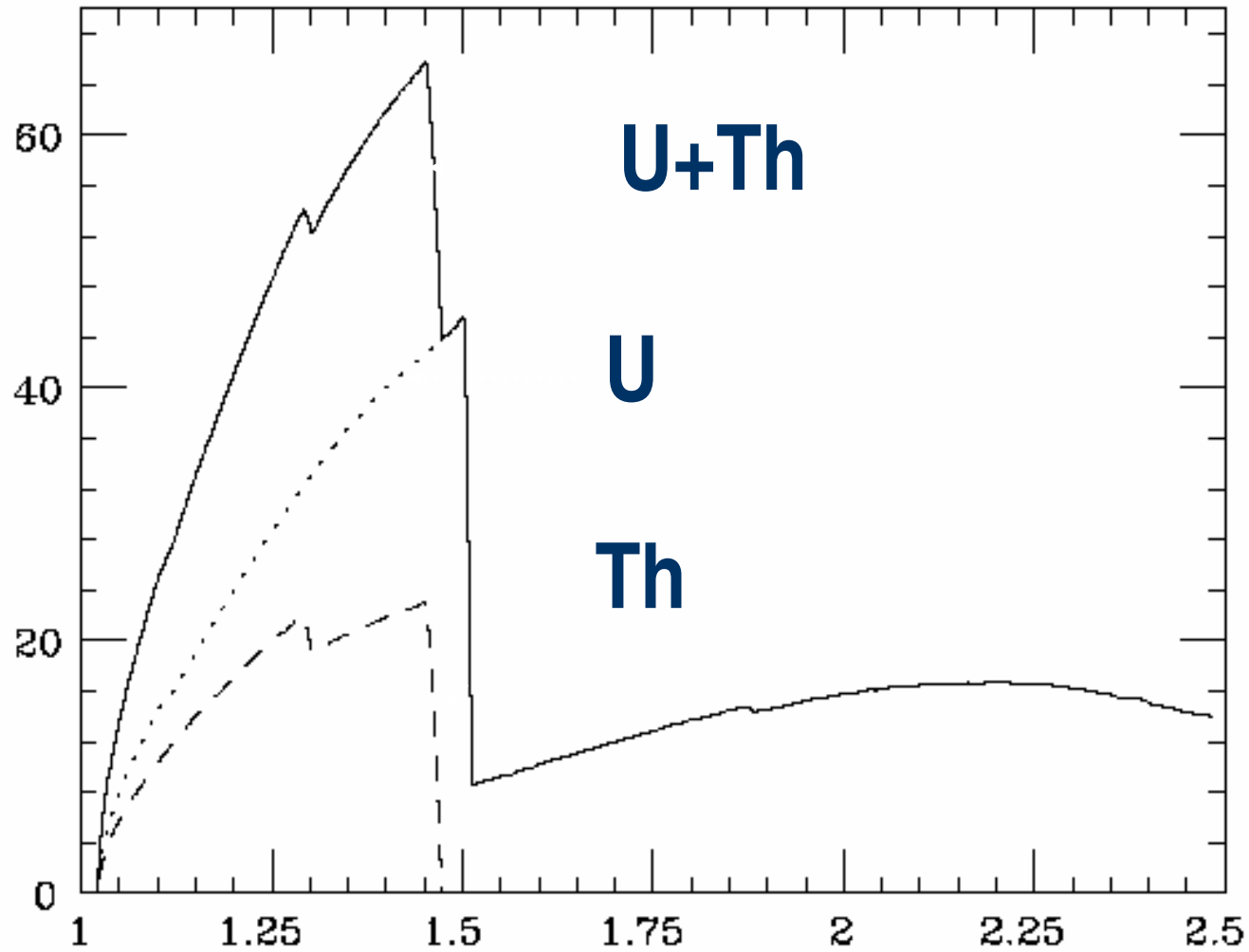
# Geo-neutrino spectra

Neutrino/decay/MeV

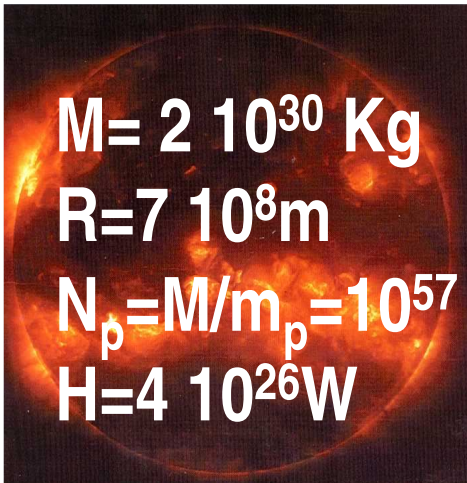


Thres

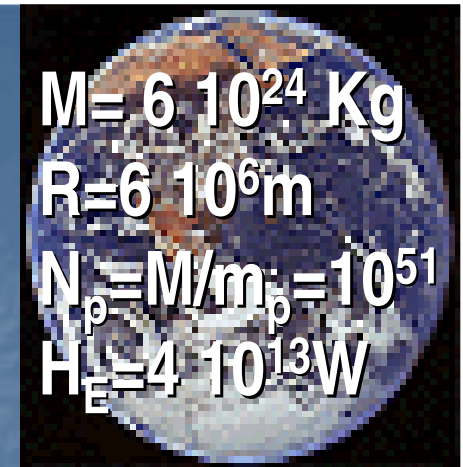
# Geo-neutrino event spectrum at Kamioka







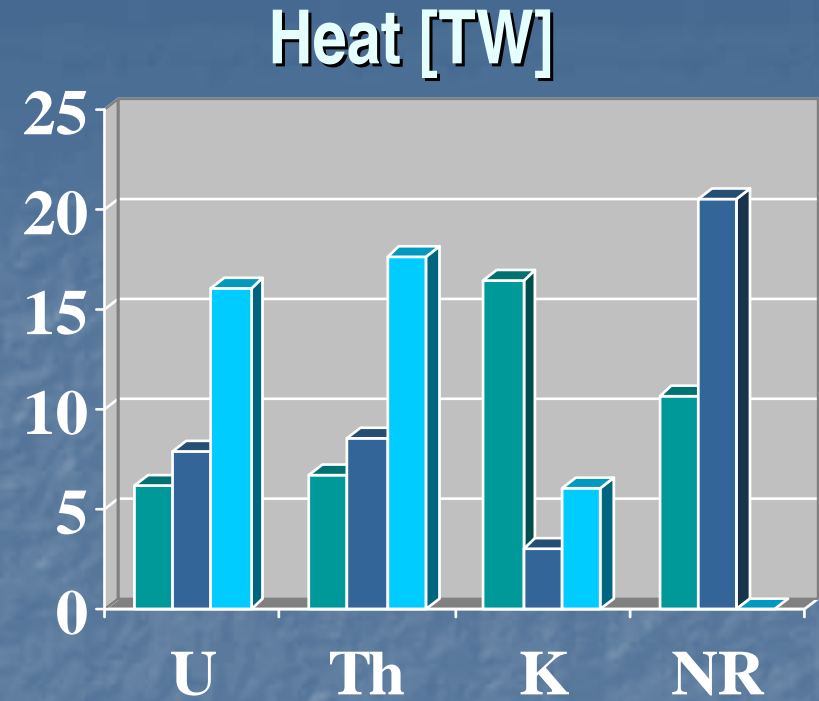
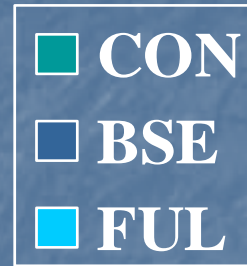
# Sun and Earth energy inventory



- The present heat flow  $H$  can be sustained by an energy source  $U$  for an age  $t$  provided that  $U > Ht$  :

	Sun [yr]	Earth [yr]
• a) chemistry: $U \approx (0.1 \text{ eV}) N_p$	$\rightarrow t_{ch} = 2 \cdot 10^3$	$t_{ch} = 5 \cdot 10^{10}$
• b) gravitation $U \approx GM^2/R$	$\rightarrow t_{gr} = 3 \cdot 10^7$	$t_{gr} = 3 \cdot 10^{11}$
• c) nuclear $U \approx (1 \text{ MeV}) N_p$	$\rightarrow t_{nu} = 2 \cdot 10^{10}$	$t_{nu} = 5 \cdot 10^9$
• Only nuclear energy is important for sustaining the Solar luminosity over the sun age, $t = 4.5 \cdot 10^9 \text{ y}$ (as proven by Gallium solar neutrino experiments).		
• All energy sources seem capable to sustain $H_E$ on geological times.		

# Reasonable models for radiogenic heat production



• A naïve chondritic model easily accounts for 3/4 of  $H_E$ , mainly from  $^{40}\text{K}$ , however  $^{40}\text{K}/\text{U}=7$ .

• In The “standard” BSE model ( $^{40}\text{K}/\text{U}=1$ ) radiogenic production is 1/2  $H_E$ , mainly from U and Th. Predictions fixed to (10-15)%

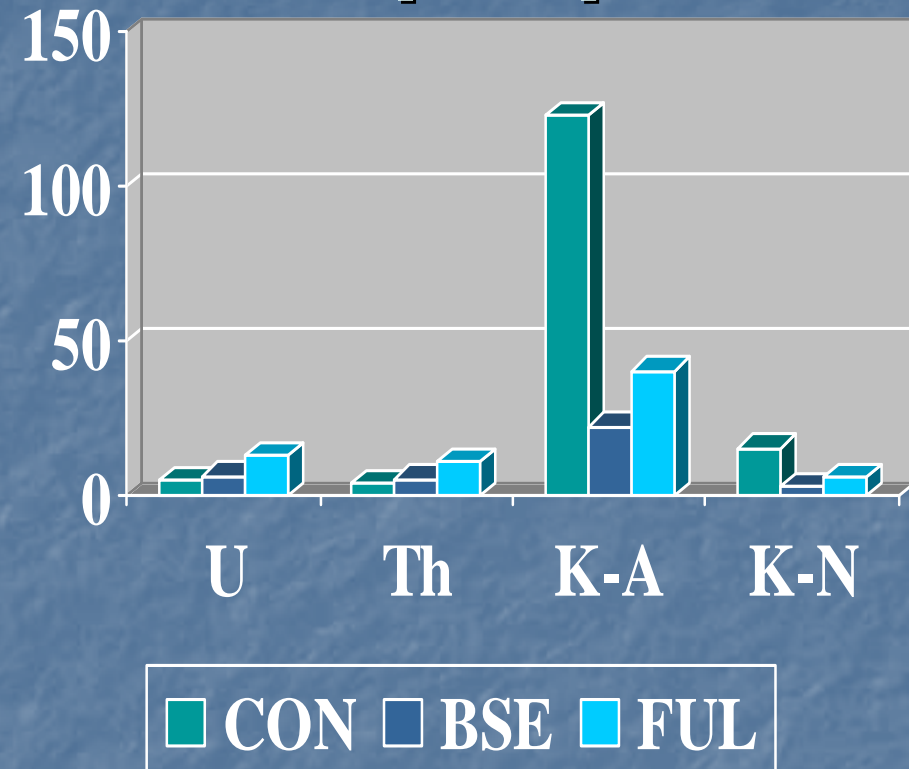
• A fully radiogenic model (imposing  $H_{\text{rad}}=40$  TW,  $\text{Th}/\text{U}=4$  and  $^{40}\text{K}/\text{U}=1$ ) is not excluded by data.

\* NR= Non Radiogenic heat

# The range of $\nu$ luminosities

- In any model, anti- $\nu$  production is dominated by  $^{40}\text{K}$ .
- Th and U anti-neutrino luminosities are in the range  $(10-20)10^{24}/\text{s}$ .

neutrino Luminosity  
[ $10^{24}/\text{s}$ ]



KA=Potassium antiv  
KN= Potassium  $\nu$

# From luminosity to fluxes and events

- Since Earth surface is  $S \approx 6 \cdot 10^{18} \text{ cm}^2$  (anti) neutrino fluxes are in the range:

$$\Phi \approx L/S \approx 10^6 \text{ cm}^{-2} \text{ s}^{-1}$$

in the same range of solar B-neutrinos\*.

- For calculating the (angle integrated) flux at a specific site one needs to know the total amounts of radioactive nuclei and their distribution\*\*:

$$\Phi = \frac{1}{4\pi} \int d^3r \frac{A(r)}{|R-r|^2} P_{ee}$$

- Estimates of the crust contribution can be provided by using geological maps of Earth crust (which distinguish CC from OC and also distinguish several layers in the CC).
- Keep in mind that for the mantle, only very rough information is available.

\*) This is different from the normal flux, which for spherical symmetry is anyhow

$$\Phi_{\nu} = L_{\nu} / 4\pi R^2$$

# Antineutrinos from the earth: a reference model and its uncertainties

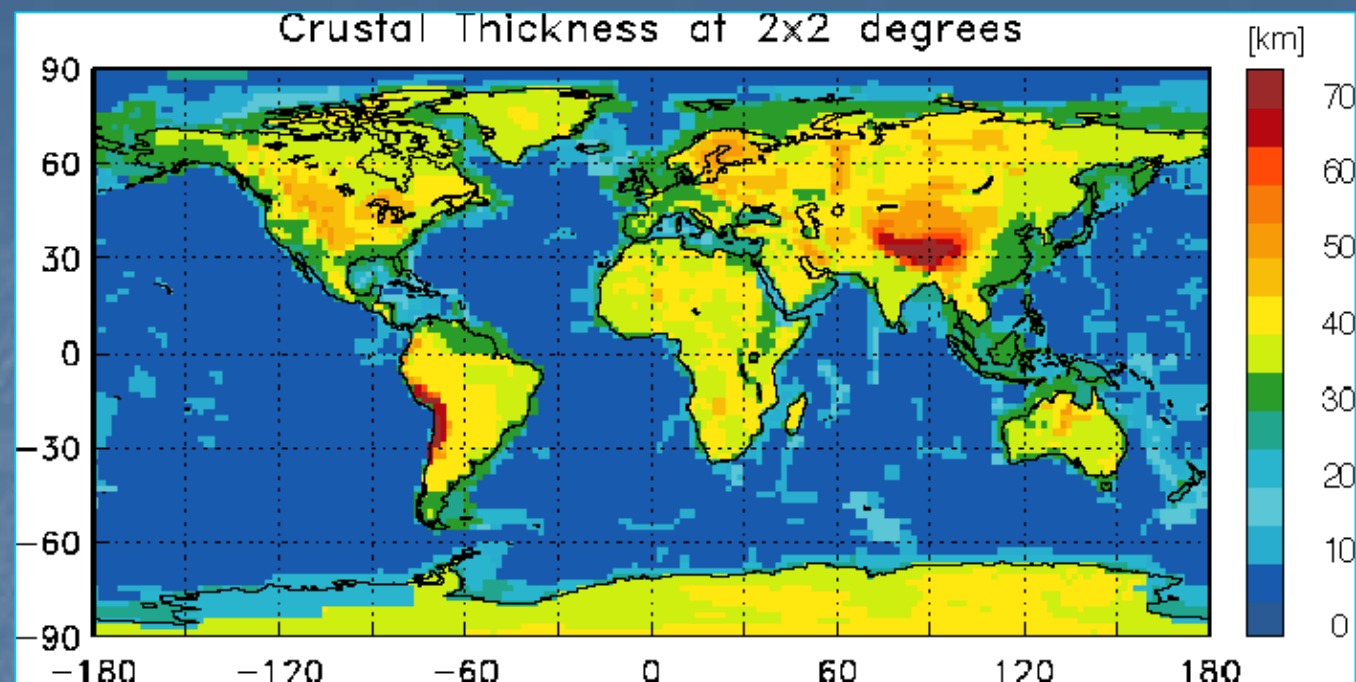
## Abstract

We predict geoneutrino fluxes in a reference model based on a detailed description of the earth's crust and mantle and using the best available information on the abundances of uranium, thorium and potassium inside earth's layers. We estimate uncertainties of fluxes corresponding to uncertainties of the element abundances. In addition to distance integrated fluxes, we also provide the differential fluxes as a function of distance from several sites of experimental interest. Event yields at several locations are estimated and their dependence on the neutrino oscillation parameters is discussed. At Kamioka we predict  $N(U + Th) = 35 \pm 6$  events for  $10^{32}$  proton · yr and 100% efficiency assuming  $\sin^2(2\theta) = 0.863$  and  $\delta m^2 = 7.3 \cdot 10^{-5} \text{eV}^2$ . The maximal prediction is 55 events, obtained in a model with fully radiogenic production of the terrestrial heat flow.

F. Mantovani et al.

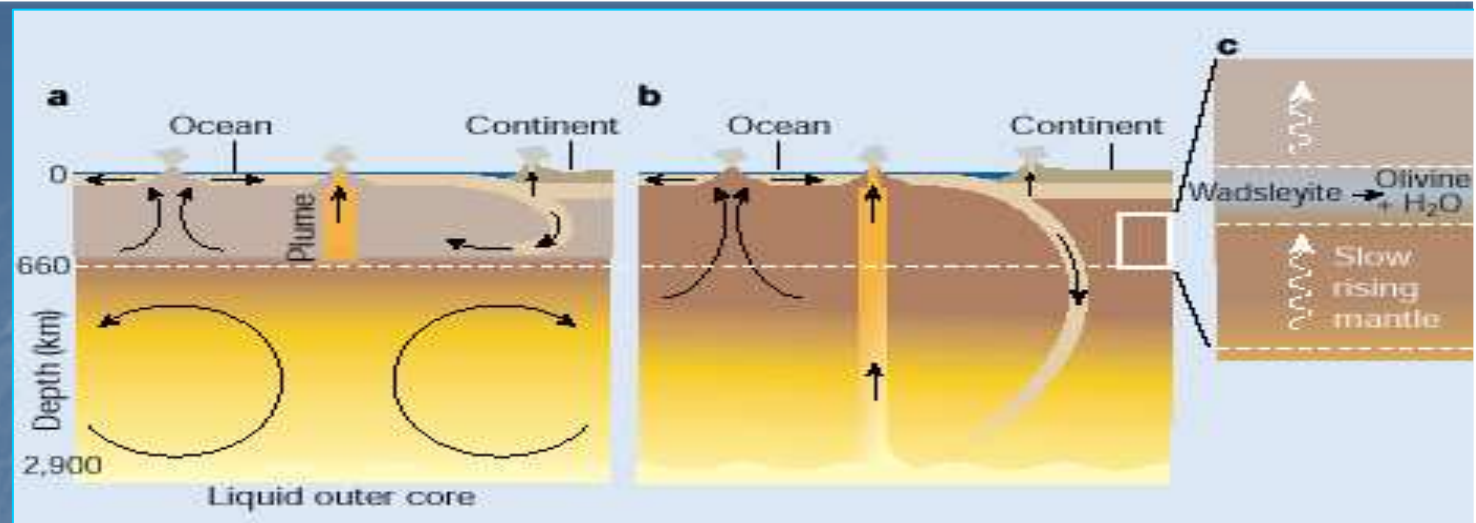
arXiv: hep- ph/ 0309013 1 Sep 2003, Phys. ReV. D

## The crust



- We use a 2x2 degrees crust map, which distinguishes several components (CC and OC, upper, middle, lower...)
- For each layer we use the average of reported values in literature for U, Th and K abundances.
- We deduce uncertainties from spread in reported data

# The mantle



- Geochemists prefer a layered mantle, however seismology prefers a wholly mixed mantle
- In the reference model we use geochemist description: for UM we take observed values and for LM the complement to BSE estimate
- We checked that that uniform mantle gives essentially the same flux (a part from sites where mantle contribution is dominant)

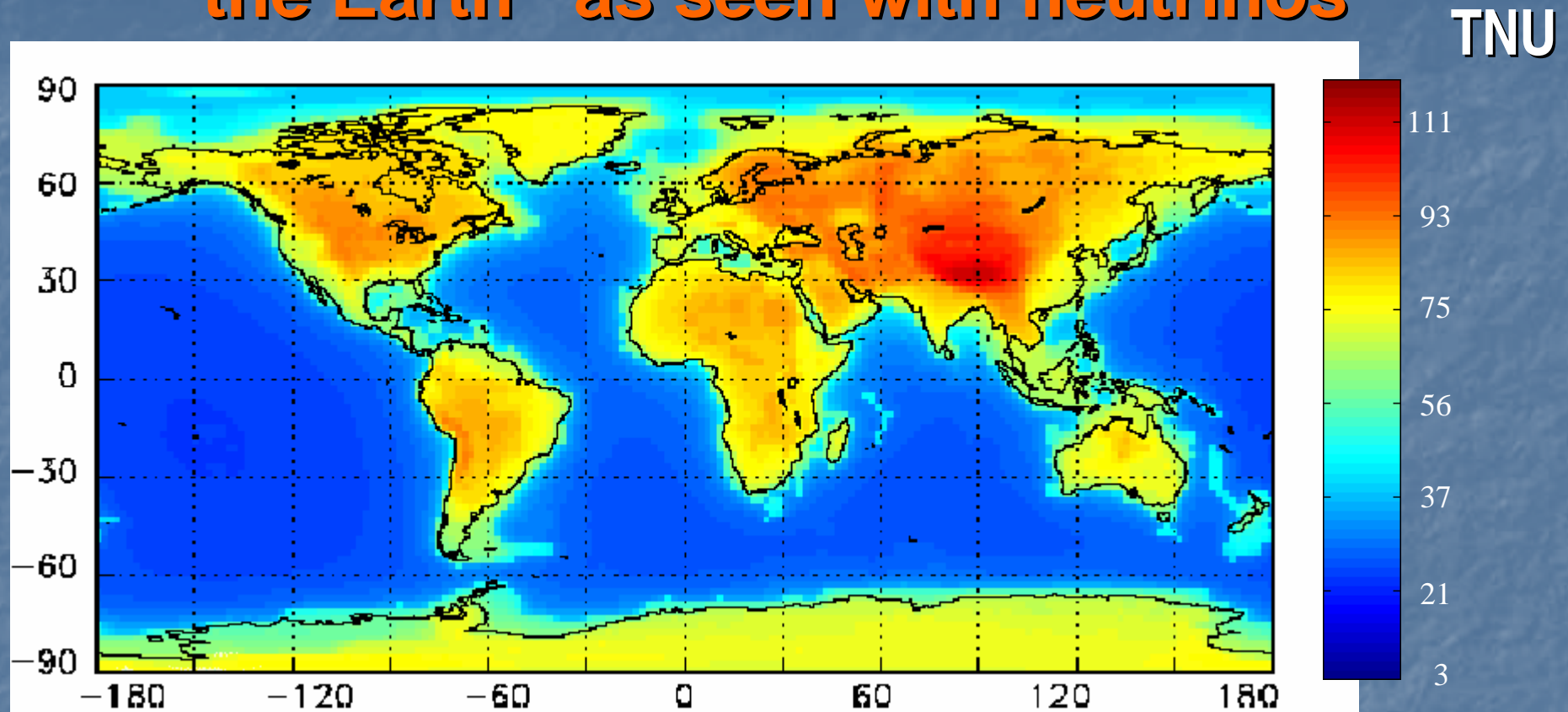
# Geo-neutrinos, Mantle Circulation and Silicate Earth

- We have studied geo-neutrino production for different models of matter circulation and composition in the mantle.
- **By using global mass balance for the Bulk Silicate Earth**, the predicted flux contribution from distant sources in the crust and in the mantle is fixed within  $\pm 15\%$  (full range).
- A detailed geological and geochemical investigation of the region near the detector has to be performed, for reducing the uncertainty from fluctuations of the local abundances to the level of the global geochemical error.
- A five-kton detector operating over four years at a site relatively far from nuclear power plants can measure the geo-neutrino signal with 5% accuracy
- **It will provide a crucial test of the Bulk Silicate Earth and a direct estimate of the radiogenic contribution to terrestrial heat.**

F Mantovani et al Hep-ph/0401085, JHEP



# The reference model : the Earth as seen with neutrinos



Predicted U+Th geoneutrino events

[1 TNU= 1 event / (  $10^{32}$  protons x year ) ]

\*look at [www.fe.infn.it/~fiorenti](http://www.fe.infn.it/~fiorenti)

# Predictions of the reference model

TABLE XII: **Total yields.**  $N_{no}$  is the total number of geoevents (U + Th) in the absence of oscillations predicted from the reference model for  $10^{32}$  proton  $\cdot$  yr (or in TNU) and  $\Delta N_{no}$  is the “ $1\sigma$ ” error.  $N_{no}^{low}$  ( $N_{no}^{high}$ ) is the minimal (maximal) prediction <sup>a</sup>.

Location	$N_{no}$	$\Delta N_{no}$	$N_{no}^{low}$	$N_{no}^{high}$
Baksan	91	13	51	131
Hawaii	22	6	10	49
Himalaya	112	15	63	154
Homestake	91	13	51	130
Kamioka	61	10	33	96
La Palma	37	8	19	67
LGS	71	11	39	106
Pyhasalmi	92	13	51	131
Sudbury	87	13	48	125
Yucca Mountain	70	11	38	106

<sup>a</sup>For  $\delta m^2 > 4 \cdot 10^{-5}$  eV<sup>2</sup> the geoevent yield is  $N = N_{\text{no}} \cdot [1 - 0.5 \sin^2(2\theta)]$ .

# Our first predictions for Kamland

(nucl-ex/0212008 ,8 December 2002. Phys Lett. 2003)

Model	N(U+Th)*
Chondritic	2.6
BSE	3.1
Fully rad.	5.1

- “The determination of the radiogenic component of terrestrial heat is an important and so far unanswered question.... ,the first fruit we can get from neutrinos, and **Kamland will get the firstlings very soon**”.

- \*Events normalized to  $0.14 \cdot 10^{32}$  p·yr,  $\epsilon=78\%$  and  $P_{ee}=0.55$ .

# Prospects for measuring terrestrial heat with geoneutrinos

- Remind that main uncertainties concern mantle.
- Consider Uranium geoneutrinos as a mean to determine U in the mantle, and thus to determine  $H_U$ , expected between 6 and 16 TW.

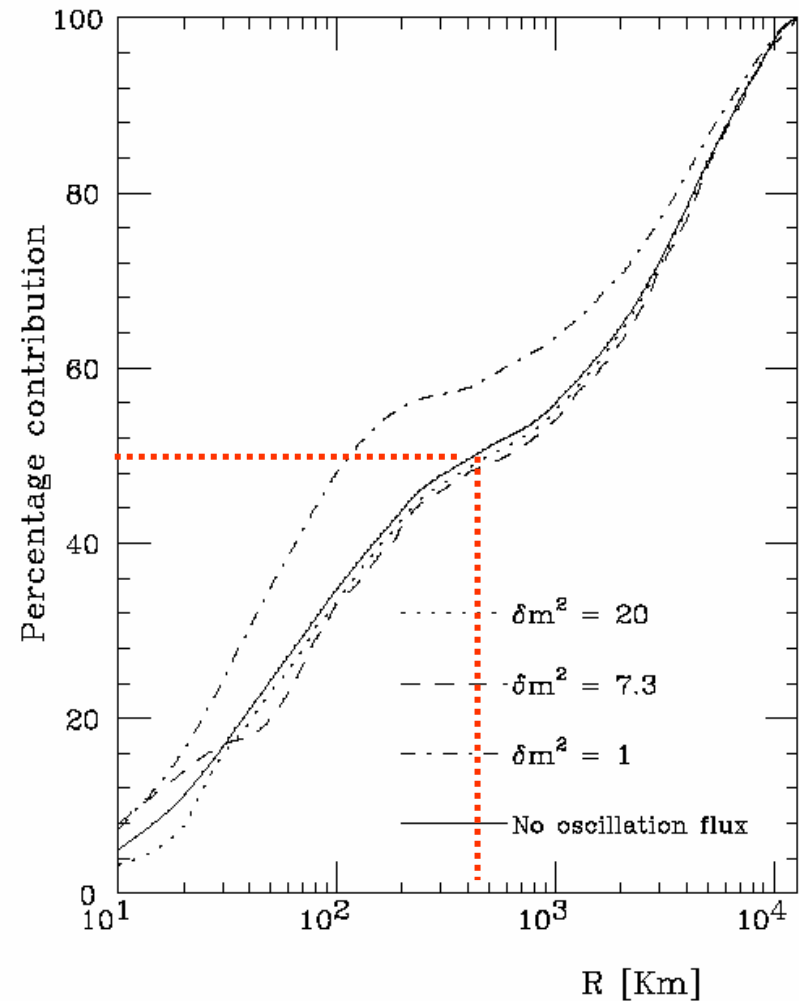
Location	Kamioka	Gran Sasso	Kamioka	Gran Sasso
Exposure** (p yr) $10^{32}$		$10^{32}$	$10^{33}$	$10^{33}$
Reactor events 207***		35	2070	350
$N_U$	18-44	20-48	180-440	200-480
$\Delta H_U$ (TW)	7.7	4.6	2.7	2.1

- An accuracy  $\Delta H(U)=(2-3)TW$  can be reached with an exposure  $10^{(32-33)}p\cdot yr^*$ .
- Larger exposure not really useful, due to uncertainties on U in the crust.

\*)1kton mineral oil =  $0.810^{32}$  p; \*\*assume 100% eff. ; \*\*\* maybe less if some reactor is switched off

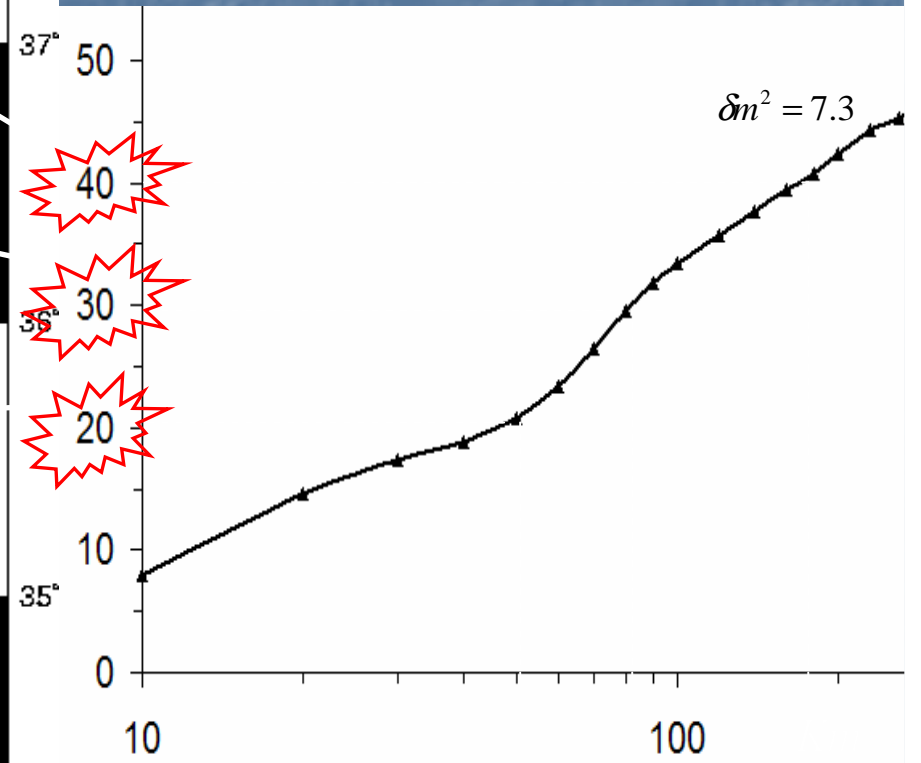
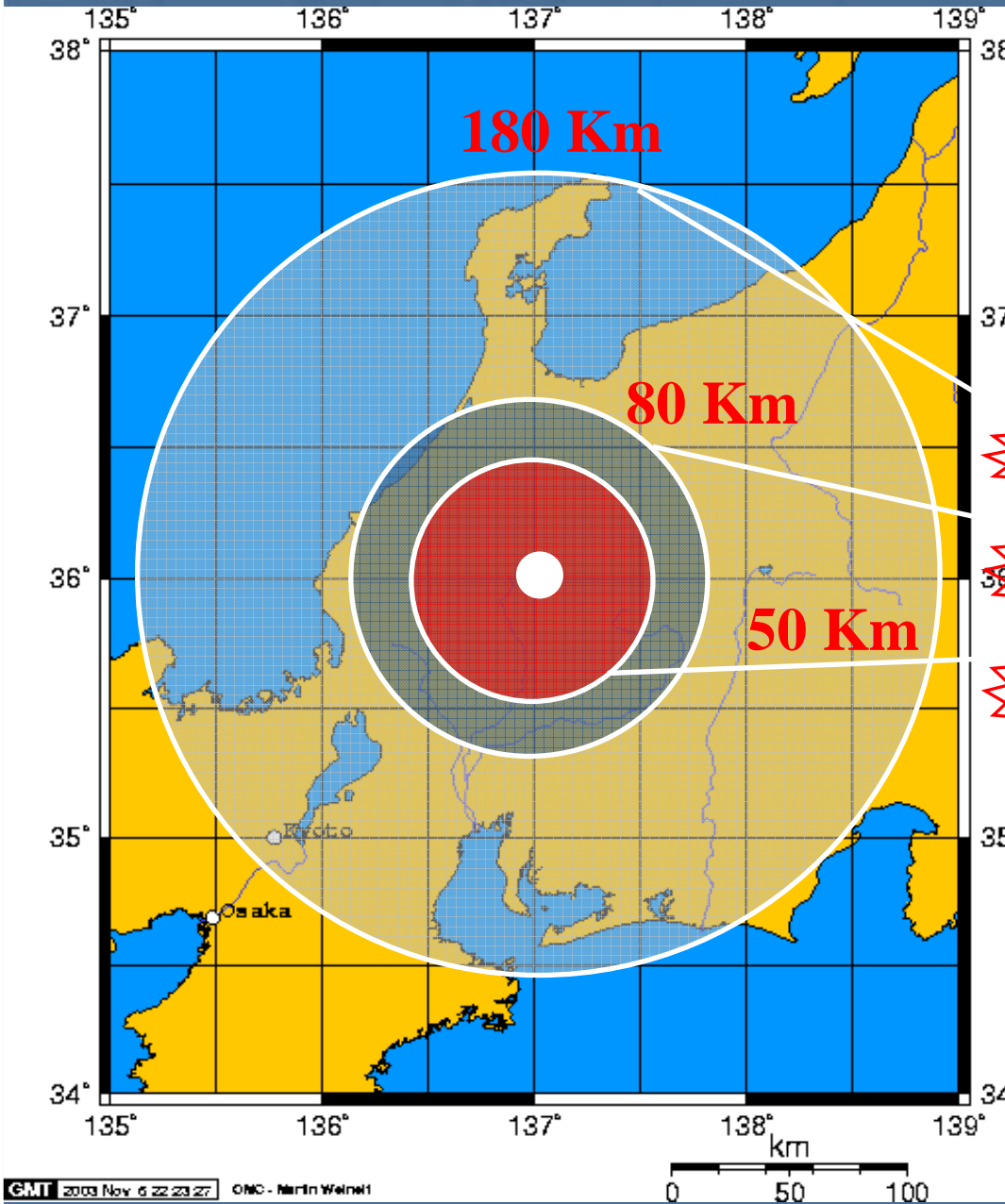
# The role of distance

- We provide estimate of the contributed flux at Kamioka as a function of distance



- We find that 50% of the flux is generated from distances larger than 400 km.

# The importance of the local contribution



Percentage contribution to the yield

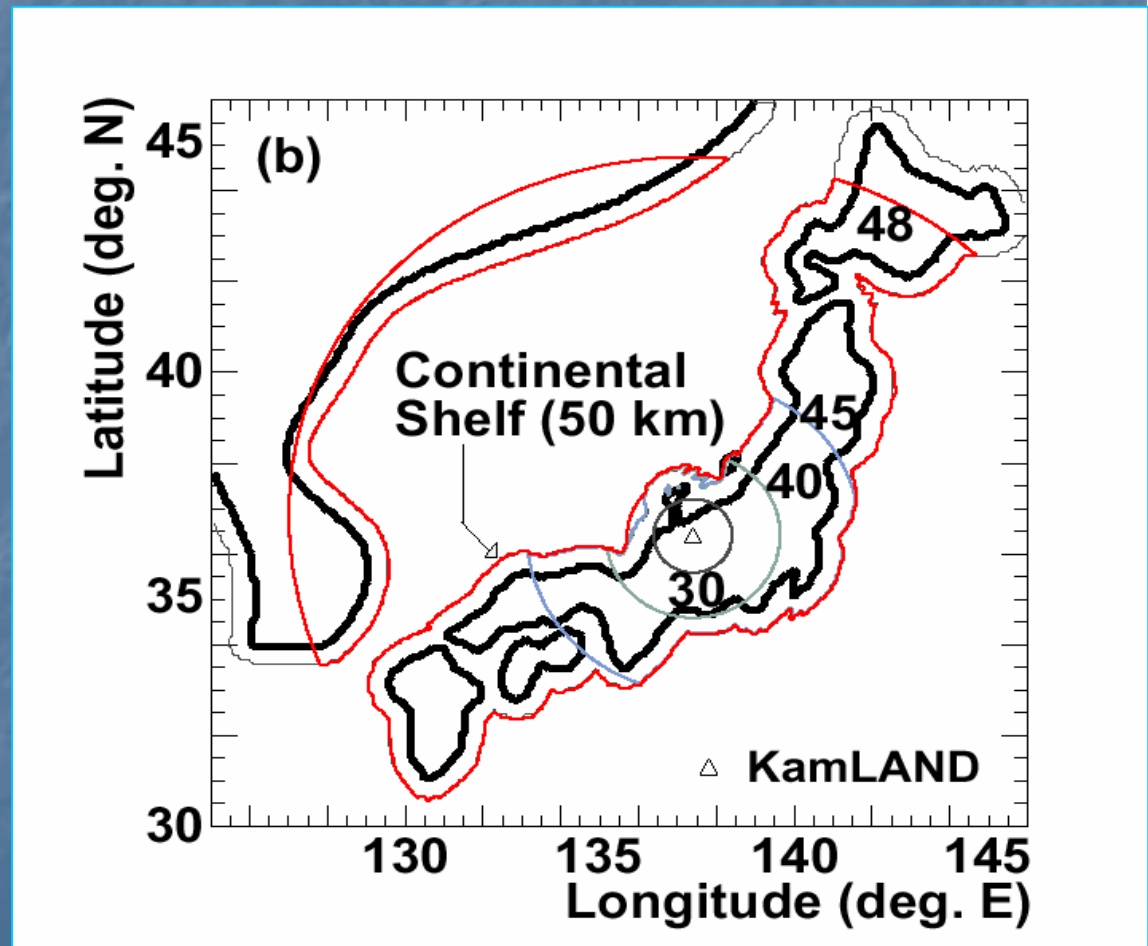
# Comparison with analysis by Mitsui\*

Best estimate by Mitsui:

$N(\text{U}+\text{Th})=32$  events  
(for  $10^{32}$  p yr)

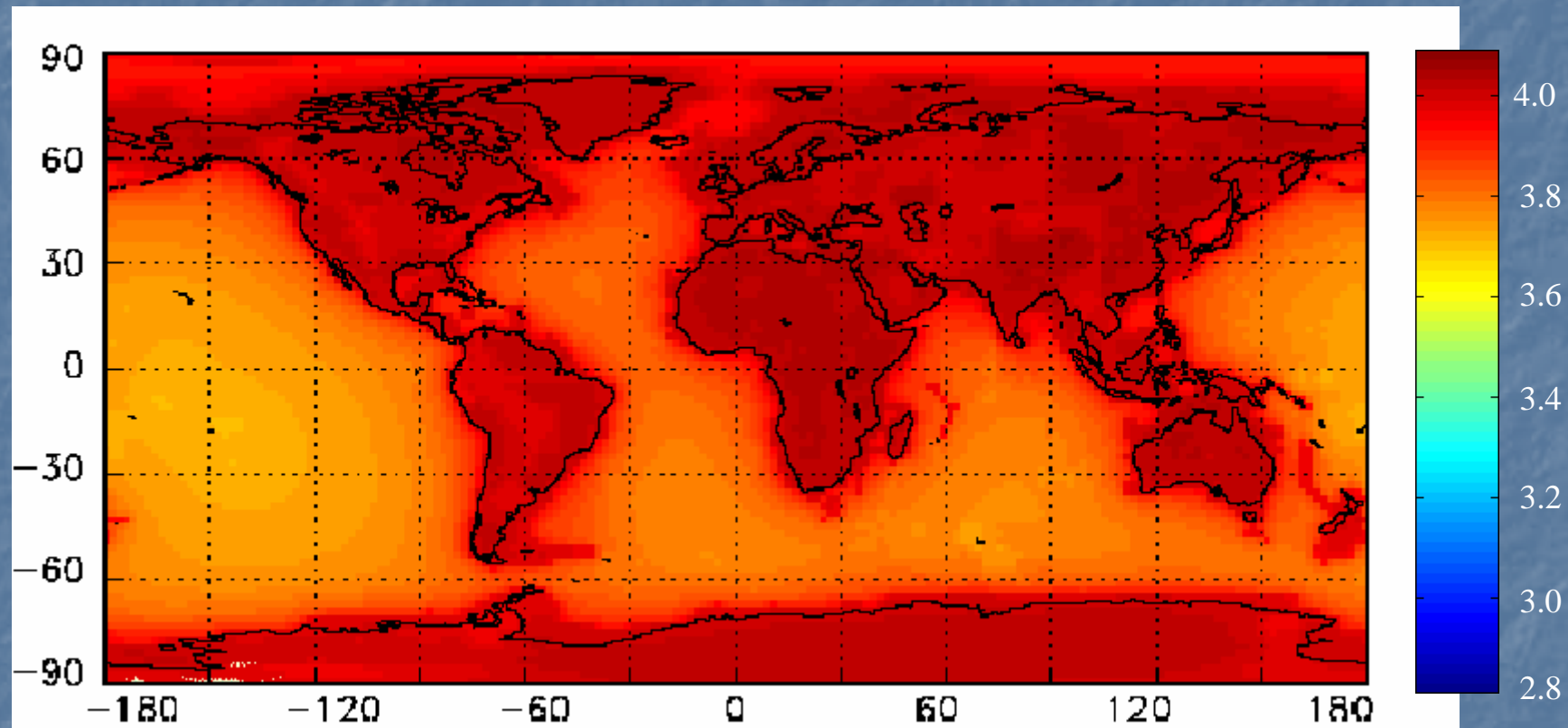
to be compared with our  
 $N(\text{U}+\text{Th})=36$

The percentage crust  
contribution is also  
estimated.



\*ICRC-2003

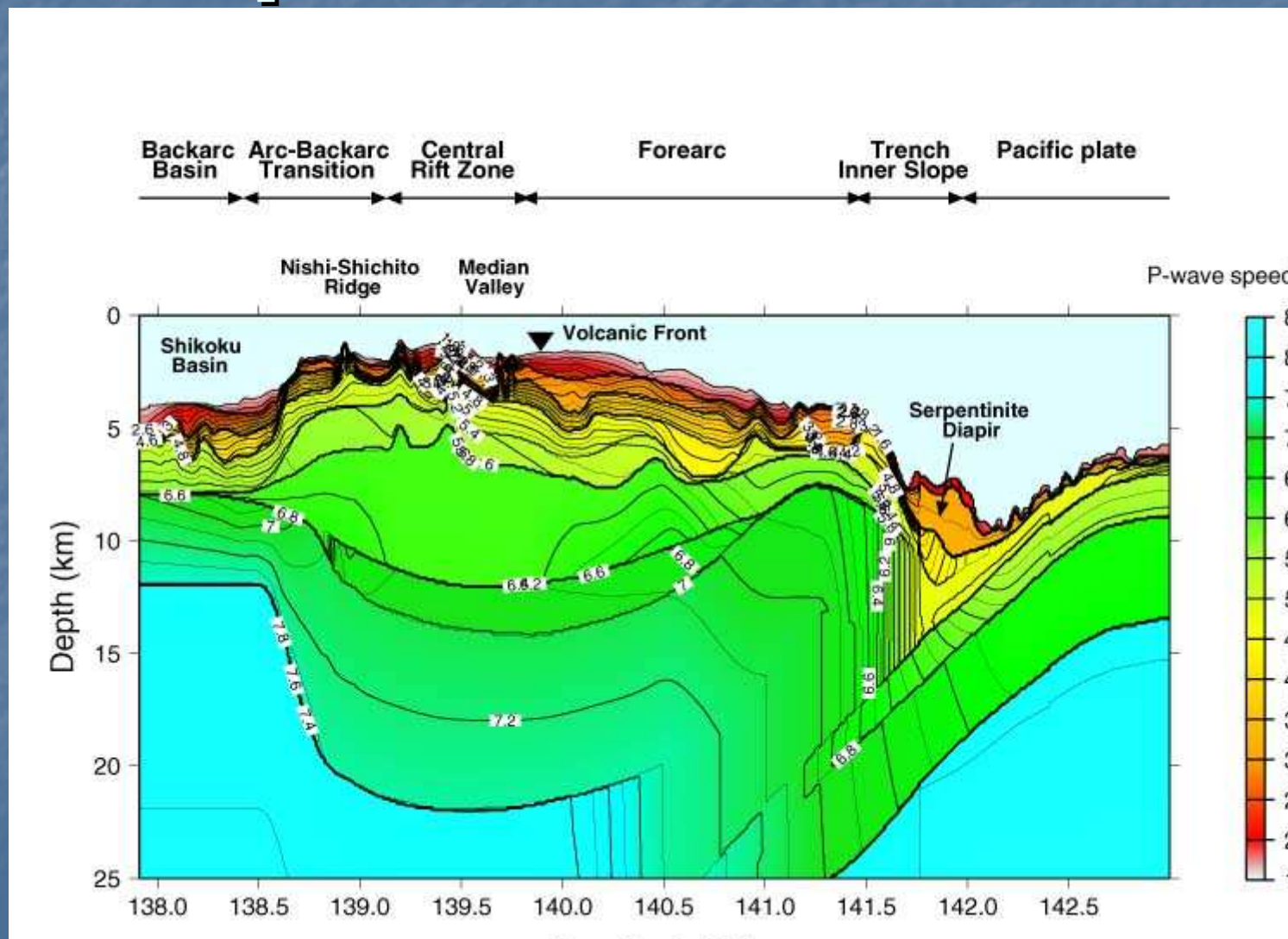
# Mappa Th/U



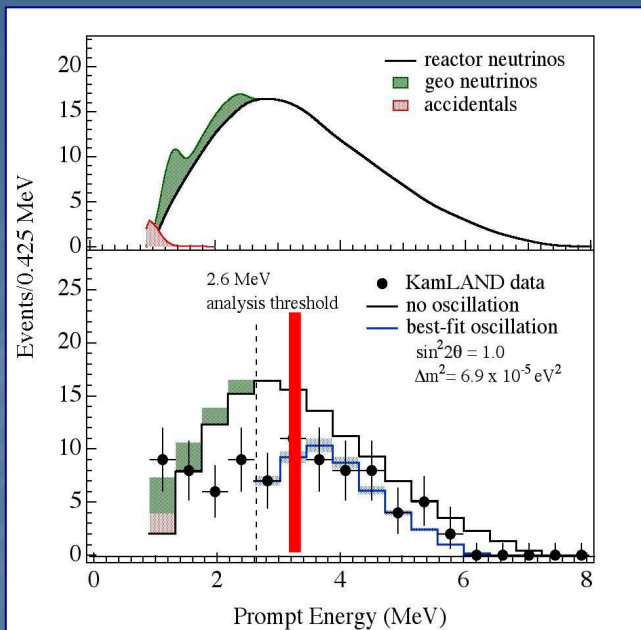
Ottenuta con: Crosta Standard + Mantello Sup. Standard:  $a(\text{U}) = 6.5 \cdot 10^9$  ;  $a(\text{Th}) = 17.3 \cdot 10^9$   
Mantello Inf. Impoverito:  $a(\text{U}) = 13.2 \cdot 10^9$  ;  $a(\text{Th}) = 52 \cdot 10^9$



# Japan cross section



# FAQ: Can we learn on neutrinos from geoneutrinos?

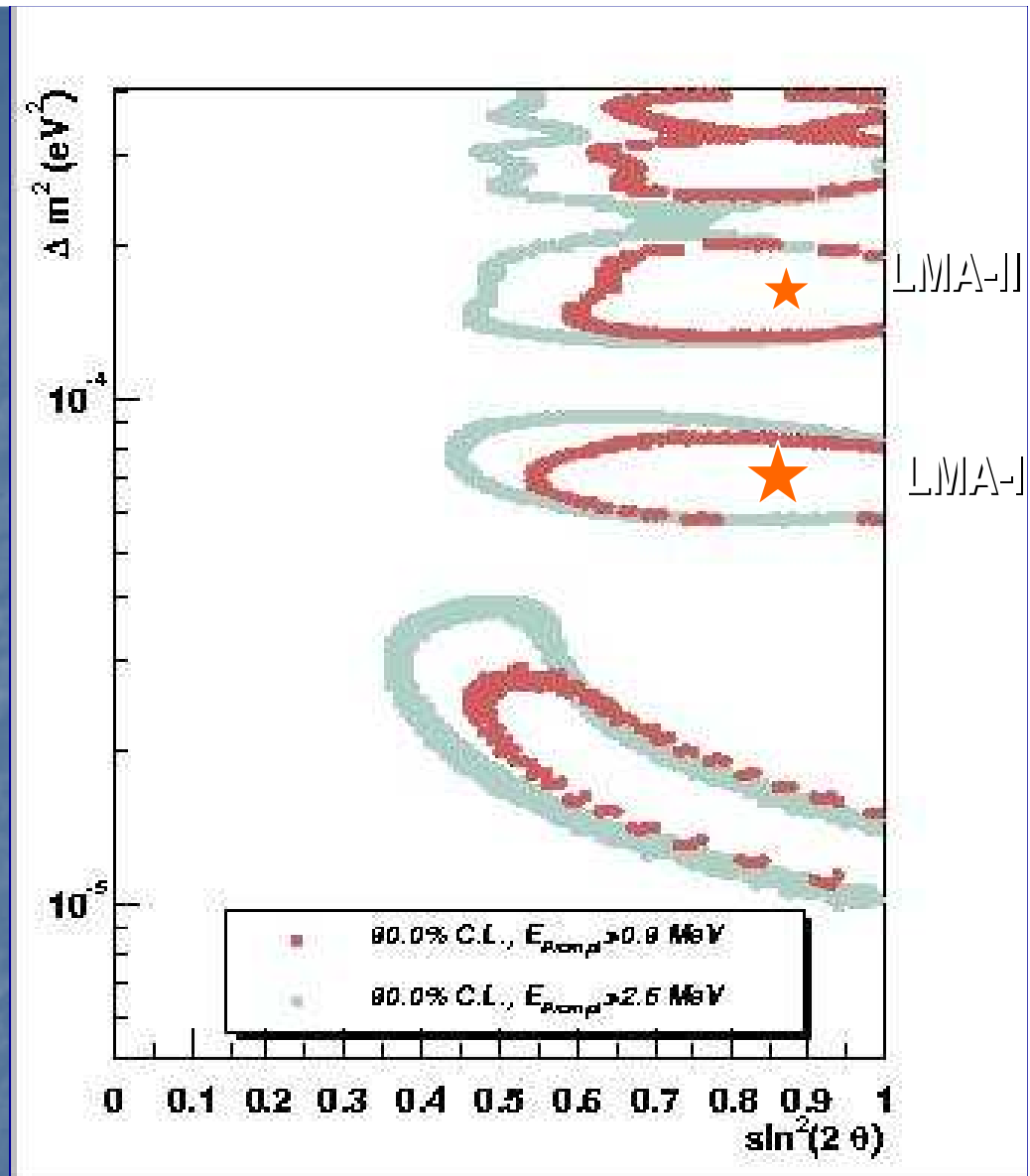


- Q1: Can you tell me U and/or Th fluxes so that I can improve determination of  $\theta$  and  $\delta m^2$ ?
- A1: No. Better you give me  $\theta$  and  $\delta m^2$  and we deduce amounts of U and Th inside the Earth

- Q2: Should I ignore anything below the red line for determining  $\theta$  and  $\delta m^2$ ?
- A2: No. I can tell you the event ratio:
  - $N(\text{Th})/N(\text{U})=0.25 \pm 0.05$
  - This follows from  $\text{Th}/\text{U}=4$ , well fixed in the solar system (meteorites, Venus, Moon) and also in the Earth.
  - This can be used to constrain mixing parameters

# We can learn on neutrinos from geoneutrinos

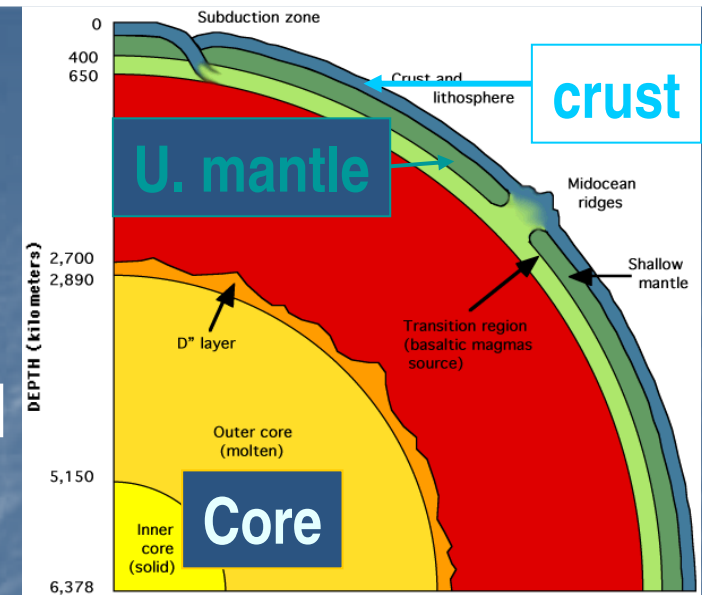
- The Th/U constraint  $N(\text{Th})/N(\text{U})=0.25\pm 0.05$  already gives some information:
  - a (very) slight preference for LMA-I
  - some reduction of the allowed parameter space.



This constraint should be kept in mind when higher statistics becomes available.

# What about Potassium?

- Earth looks depleted by a factor of seven with respect to oldest meteorites.
- Elements as heavy as Potassium should not have escaped from a planet as big as Earth.
- Most reasonable assumption is that it volatilized in the formation of planetesimals from which Earth has accreted (heterogeneous accretion),
- However, at high pressure Potassium behaves as a metal and thus it might have been buried in the Earth core, where it could provide the energy source of the terrestrial magnetic field (see eg. Rama Nature 2003).
- A long standing debate...



# Potassium in the core?

W F Mc Donough “Compositional Model for the Earth’s Core”  
2003

- “Potassium is commonly invoked as being sequestered into the Earth’s core due to:
  - (i)potassium sulfide found in some meteorites;
  - (ii)effects of high-pressures–d-electronic transitions;
  - (iii)solubility of potassium inFe–S (and Fe–S–O)liquids at high pressure.
- **Each of these is considered below and rejected....”**

# Uranium and Th in the core?

W F Mc Donough “Compositional Model for the Earth’s Core”  
(2003)

- “An Earth’s core containing a significant amount of radioactive elements has been proposed by Herndon (1996).
- This model envisages a highly reduced composition for the whole Earth and, in particular, for the core.
- **Unfortunately, Herndon has developed a core compositional model that is inconsistent with chemical and isotopic observations of the Earth’s mantle and a chondritic planetary composition...**”

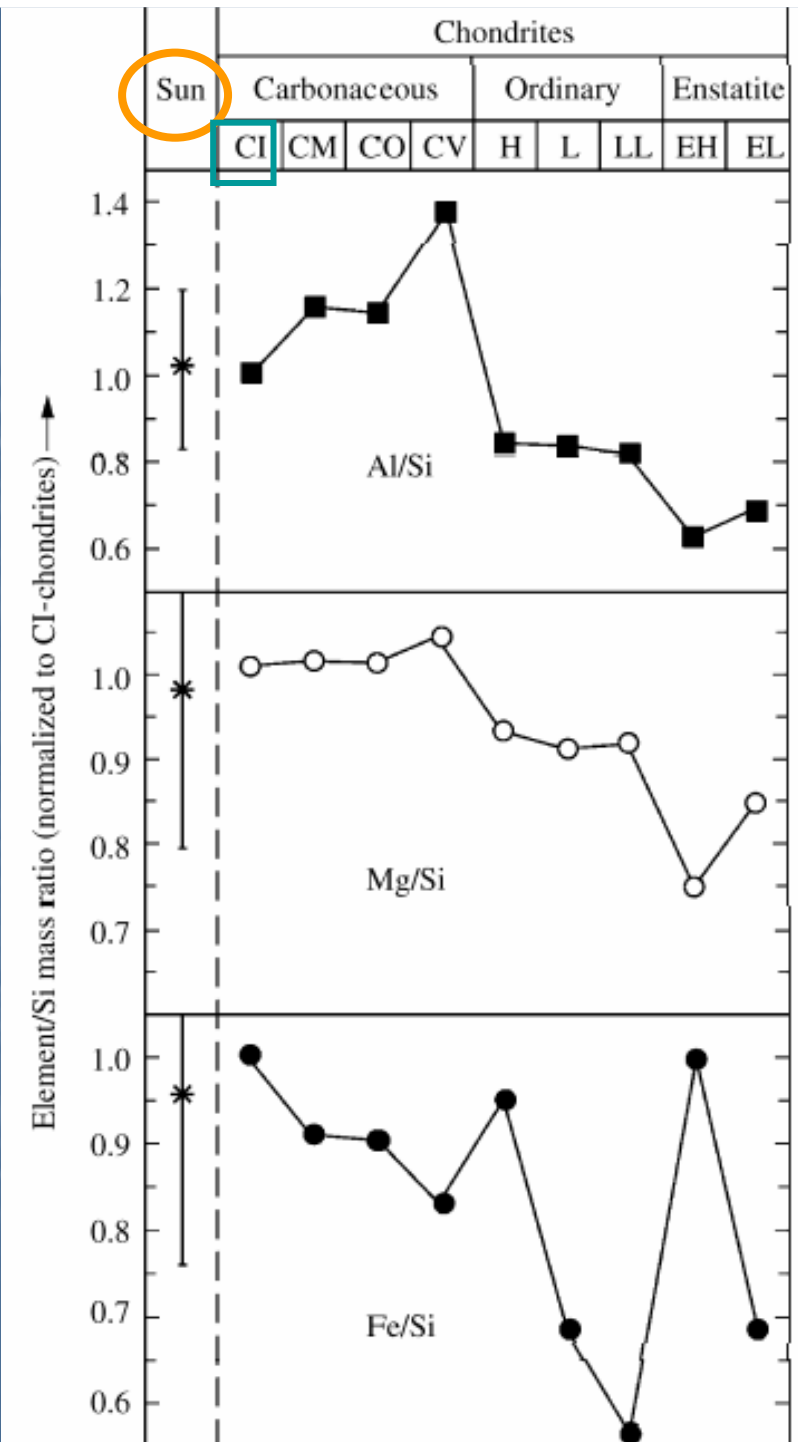
# Earth's and meteorites (Theory)



- Since we do not have access to most of the Earth's interior, Earth's composition is estimated from the similarity with CI- Chondritic meteorites
- Composition is not the same due to the fact that “volatile” elements have escaped during Earth's formation.
- One identifies in the present Earth' “refractory& lithophile” elements, (e.g Al) which should be kept without loss and have not fallen down to the Earth's core
- One then rescales abundances of other elements to those of meteorites.
-

# Why CI-chondrites ?

- The best match between solar abundances and meteoritic abundances is with CI-meteorites.
- H.Palme and Hugh St. C.O'Neill “Cosmochemical Estimates of Mantle Composition” (2003)





# The chondritic nature of the Mantle

**Table 1** Composition of the mantle of the Earth assuming average solar system element ratios for the whole Earth.

	<i>Earth's mantle solar model</i>	<i>Earth's mantle based on composition of upper mantle rocks<sup>a</sup></i>
MgO	35.8	36.77
SiO <sub>2</sub>	51.2	45.40
FeO	6.3	8.10
Al <sub>2</sub> O <sub>3</sub>	3.7	4.49
CaO	3.0	3.65

- “The remarkable result of this exercise is that by assuming solar element abundances of rock-forming elements in the bulk Earth leads to a mantle composition that is in basic agreement with the mantle composition derived from upper mantle rocks”.

H.Palme and Hugh St. C.O'Neill “Cosmochemical Estimates of Mantle Composition” (2003)

# CI or EH ?

- A. M. Hofmeister and R.E. Criss (2003, under review): **Earth's Heat Flux Revised and Linked to Chemistry:**
- "...The Earth and the Moon share a common oxygen isotope ratio with the enstatite chondrites (classes EH and EL) and enstatite achondrites (aubrites) that is distinct from that of CI chondrites and all other meteorite classes (Clayton, 1993), necessitating large contributions from the enstatite chondrites in Earth models (Javoy, 1995; Lodders, 2000).
- In addition, the Earth is ~30 wt % Fe, almost all of which resides in its massive core. Earth's precursor materials must supply this amount. The CI model falls short, providing only 18 wt % Fe, but the EH type has just the right amount..."

# Herndon, PNAS2003

## Nuclear georeactor origin of oceanic basalt $^3\text{He}/^4\text{He}$ , evidence, and implications

J. Marvin Herndon\*

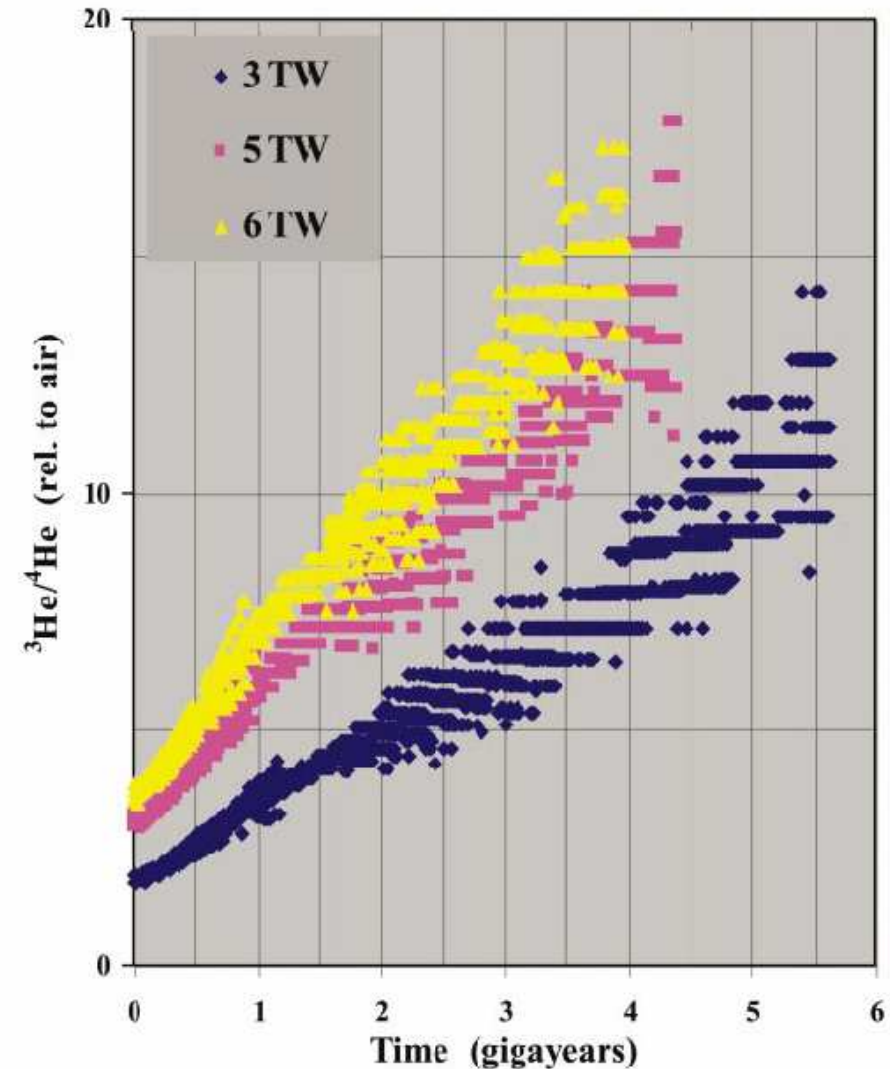
Nuclear georeactor numerical simulation results yield substantial  $^3\text{He}$  and  $^4\text{He}$  production and  $^3\text{He}/^4\text{He}$  ratios relative to air ( $R_A$ ) that encompass the entire 2-SD ( $2\sigma$ ) confidence level range of tabulated measured  $^3\text{He}/^4\text{He}$  ratios of basalts from along the global spreading ridge system. Georeactor-produced  $^3\text{He}/^4\text{He}$  ratios are related to the extent of actinide fuel consumption at time of production and are high near the end of the georeactor lifetime. Georeactor numerical simulation results and the observed high  $^3\text{He}/^4\text{He}$  ratios measured in Icelandic and Hawaiian oceanic basalts indicate that the demise of the georeactor is approaching. Within the present level of uncertainty, one cannot say precisely when georeactor demise will occur, whether in the next century, in a million years, or in a billion years from now.



**Few km U droplet, providing a natural breeder, releasing few TW over Earth's lifespan.**

# Can Herndon breeder reactor self sustain ?

- Assume Uranium and Thorium km-drops can accumulate in the center.
- Initial criticality was possible due to the high  $^{235}\text{U}/^{238}\text{U}$  enrichment of approximately 30 % at that time.
- Furthermore, the  $^{238}\text{U}/^{239}\text{Pu}/^{235}\text{U}$  conversion cycle guarantees also a stabilized  $^{235}\text{U}/^{238}\text{U}$  enrichment of about 10%



- 3TW can be sustained for the whole Earth's history
- 6TW would have exhausted  $^{238}\text{U}$  at 4Gyr

## Detecting a Nuclear Fission Reactor at the Center of the Earth

*R. .S. Raghavan*

*Bell Laboratories, Lucent Technologies, Murray Hill NJ*

*and*

*INFN Laboratori Nazionali del Gran Sasso, Italy*

A natural nuclear fission reactor with a power output of 3-to 10 Terawatt at the center of the earth has been proposed as the energy source of the earth's magnetic field. The proposal can be *directly* tested by a massive liquid scintillation detector that can detect the signature spectrum of antineutrinos from the geo-reactor as well as the *direction* of the antineutrino source. Such detectors are now in operation or under construction in Japan/Europe. However, the clarity of both types of measurements may be limited by background from antineutrinos from surface power reactors. Future U.S. detectors, relatively more remote from power reactors, may be more suitable for achieving unambiguous spectral and directional evidence for a 3TW geo-reactor

# Estimates of terrestrial heat flow

